

TOWARDS A MECHANISM FOR FORMATION OF SILICON CARBIDE CRYSTALS IN AGB STARS

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Silicon carbide (SiC) grains comprise a significant fraction of the dust found around carbon-rich AGB stars. Their presence in the interstellar medium is thought to originate from self-assembly of organosilicon building blocks, including previously observed species such as carborundum and cyclic silicon dicarbide (*c*-SiC₂). However, the actual formation mechanisms of even these simple silicon-bearing organic molecules remains elusive. Here it is proposed that disilyne (Si₂H₂) reacts barrierlessly with abundant acetylene (C₂H₂) on a spin-conserving potential to form C₂Si₂H₄. This species has been shown in experimental and theoretical studies^a to photoisomerize under UV irradiation resulting in the formation of several species, one being a *c*-SiC₂ precursor and another being a highly polar species capable of supporting a dipole-bound electron. This strongly dipolar C₂Si₂H₄ isomer may represent the missing link supporting the molecular aggregation hypothesis for SiC formation. Importantly, its polarity drives molecular aggregation, and, after subsequent oxidation to C₂Si₂, its heteronuclear linkages are well-prepared for SiC nucleation, presumably initiated by a shock-wave pulsation event. Past theoretical studies by our group^{b,c} are combined with new results, computed at the DFT and coupled-cluster levels of theory, to support the proposed mechanism.

^aLutz J.J., *Inorganics*, *submitted*

^bLutz J.J., Duan X.F., et al. *J. Chem. Phys.* 148, 174309 (2018)

^cByrd J.N., Lutz J.J., et al. *J. Chem. Phys.* 145, 024312 (2016)