

# ROTATIONAL SPECTROSCOPY OF THE LOW ENERGY CONFORMER OF 2-METHYLBUTYRONITRILE AND SEARCH FOR IT TOWARD SAGITTARIUS B2(N2)

HOLGER S. P. MÜLLER, NADINE WEHRES, OLIVER ZINGSHEIM, FRANK LEWEN, STEPHAN SCHLEMMER, *I. Physikalisches Institut, Universität zu Köln, Köln, Germany*; JENS-UWE GRABOW, *Institut für Physikalische Chemie und Elektrochemie, Gottfried-Wilhelm-Leibniz-Universität, Hannover, Germany*; ROBIN T. GARROD, *Departments of Chemistry and Astronomy, The University of Virginia, Charlottesville, VA, USA*; ARNAUD BELLOCHE, KARL M. MENTEN, *Millimeter- und Submillimeter-Astronomie, Max-Planck-Institut für Radioastronomie, Bonn, NRW, Germany*.

Quite recently, some of us detected *iso*-propyl cyanide as the first branched alkyl molecule in space.<sup>a</sup> The identification was made in an ALMA Cycle 0 and 1 molecular line survey of Sagittarius B2(N) at 3 mm. The branched isomer was only slightly less abundant than its straight-chain isomer with a ratio of about 2 : 5. While initial chemical models favored the branched isomer somewhat, more recent models<sup>b</sup> are able to reproduce the observed ratio. Moreover, the models predicted that among the next longer butyl cyanides (BuCNs) 2-methylbutyronitrile (2-MBN) should be more abundant than both *n*-BuCN and 3-MBN by factors of around 2, with *t*-BuCN being almost negligible.

With the rotational spectra of *t*- and *n*-BuCN studied,<sup>c</sup> we investigated those of 2-MBN and 3-MBN between ~40 and ~400 GHz by conventional absorption spectroscopy and by chirped-pulse and resonator Fourier transform microwave (FTMW) spectroscopy. The analyses were guided by quantum-chemical calculations.

Here we report the analysis of the low-energy conformer of 2-MBN and a search for it in our current ALMA data. Two additional conformers are higher by ~250 and ~280 cm<sup>-1</sup>. The low-energy conformer displays a very rich rotational spectrum because of its great asymmetry ( $\kappa \approx 0.14$ ) and large *a*- and *b*-dipole moment components. Accurate <sup>14</sup>N quadrupole coupling parameters were obtained from the FTMW spectral recordings.

<sup>a</sup>A. Belloche, R. T. Garrod, H. S. P. Müller, and K. M. Menten, *Science* **345** 1584.

<sup>b</sup>R. T. Garrod, A. Belloche, H. S. P. Müller, and K. M. Menten, *Astron. Astrophys.*, in press; doi: 10.1051/0004-6361/201630254.

<sup>c</sup>Z. Kisiel, *Chem. Phys. Lett.* **118** 134; M. H. Ordu et al., *Astron. Astrophys.* **541** A121.