## INFRARED SPECTROSCOPY OF DISILICON-CARBIDE, Si<sub>2</sub>C

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Small silicon and carbon containing molecules are thought to be important building blocks of interstellar grains. Some of them have been detected in circumstellar environments of late-type stars by means of rotational spectroscopy e.g., SiC, SiC<sub>2</sub>, Si<sub>2</sub>C, c-SiC<sub>3</sub>, SiC<sub>4</sub>, while centro-symmetric species, e.g., C<sub>3</sub>, C<sub>4</sub>, C<sub>5</sub>, Si<sub>2</sub>C<sub>2</sub>, Si<sub>2</sub>C<sub>3</sub>, can only be detected by vibrational transitions, mainly in the infrared. In view of a new generation of high resolution infrared telescope instruments, e.g., EXES (Echelon-Cross-Echelle Spectrograph) onboard SOFIA (Observatory for Infrared Astronomy) and TEXES (Texas Echelon Cross Echelle Spectrograph) at the Gemini-North observatory, accurate laboratory data of small siliconcarbides in the infrared region are of high demand. In this talk we present first laboratory data of the Si<sub>2</sub>C asymmetric stretching mode at 1200 cm<sup>-1</sup>. A pulsed Nd:YAG-laser is used to vaporize a solid target of silicon exposed to a dilute sample of methane in helium buffer gas. Si<sub>2</sub>C is formed in an adiabatic expansion of a supersonic jet and radiation of a quantum cascade laser is used to record rotationally resolved spectra. To date, 160 ro-vibrational lines and have been assigned to the asymmetric stretching vibration of Si<sub>2</sub>C, and derived molecular parameters are in excellent agreement with ab initio calculations. In our global fit analysis recently published microwave laboratory data (McCarthy *et al.* 2015)<sup>a</sup> and astronomical data (Cernicharo *et al.* 2015)<sup>b</sup> were taken into account. Our new results allow for the identification of Si<sub>2</sub>C by means of high resolution infrared astronomy towards the warm background of carbon-rich stars.

<sup>&</sup>lt;sup>a</sup>McCarthy M.C., Baraban J.H., Changala P.B., Stanton J.F., Martin-Drumel M.A, Thorwirth S., et al., *J. Chem. Phys. Lett.* **6**, 2107–2111 (2015).

<sup>&</sup>lt;sup>b</sup>Cernicharo J., McCarthy M.C., Gottlieb C.A., Agundez M., Velilla Prieto L., Baraban J.H., et al. Astrophys. J. Lett. 806,L3 (2015).