

# FAR-INFRARED SPECTROSCOPY OF SHORT-LIVED SPECIES

HIROYUKI OZEKI

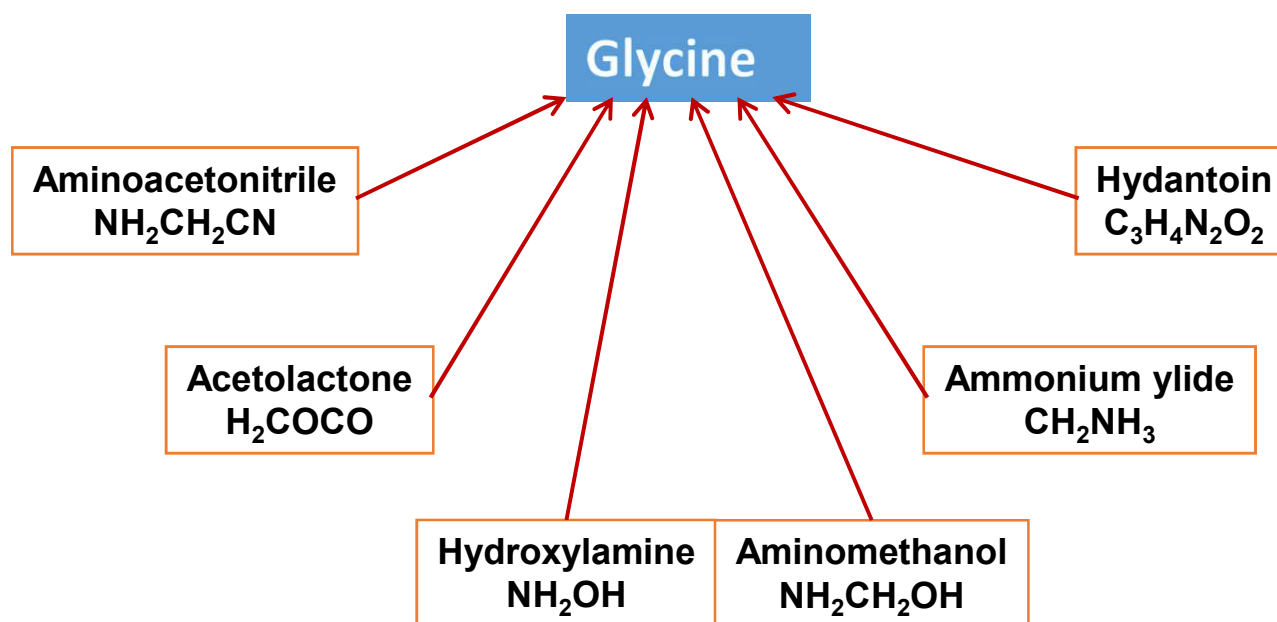
Toho University, Funabashi, JAPAN

RI05 73<sup>rd</sup> ISMS mini-Symposium Far-Infrared Spectroscopy

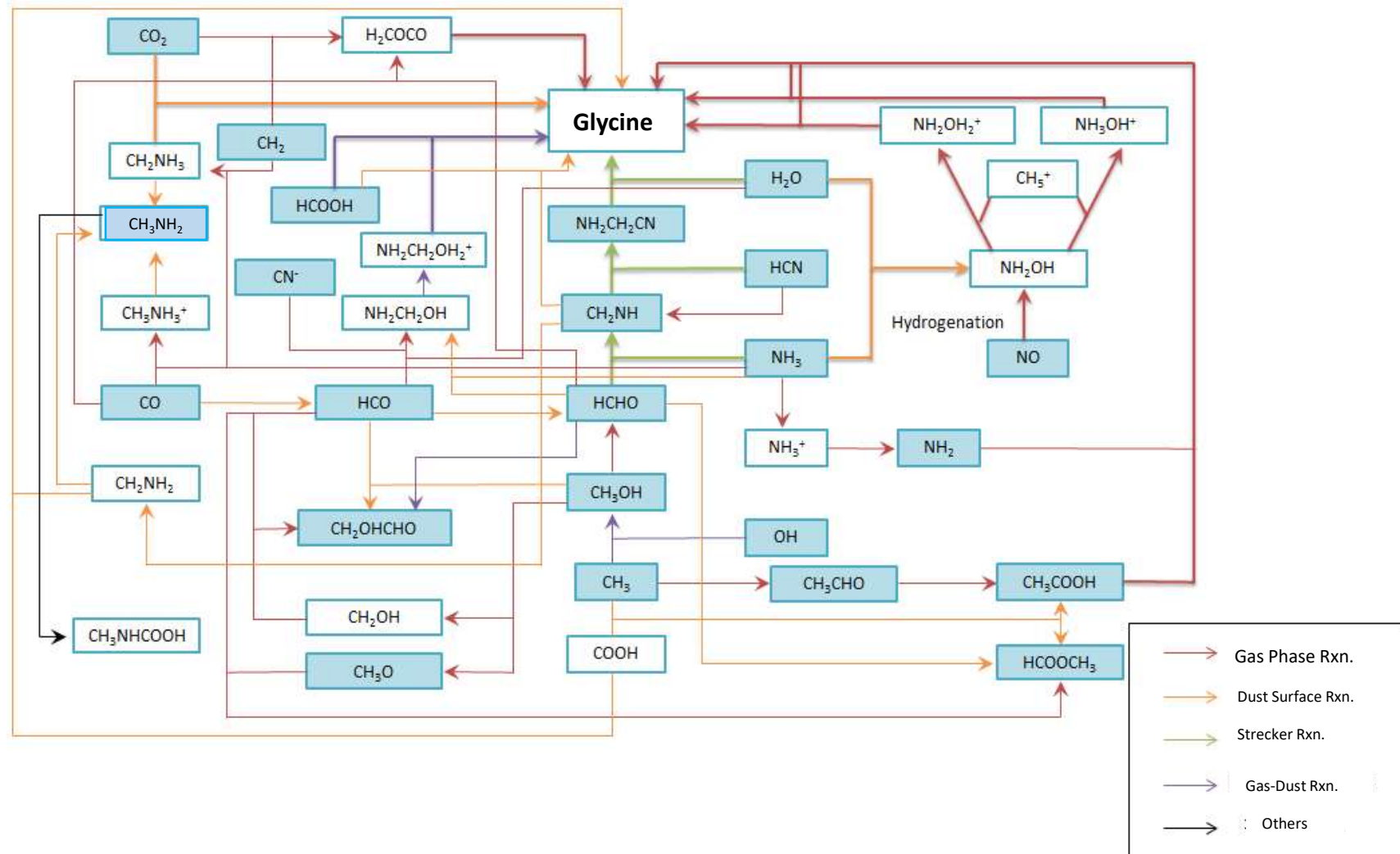
# Contents

- Spectroscopy of radicals
  - Motivation
  - “Some” keys for spectroscopy of radicals
- 3 examples
  - Methylene ( $\text{CD}_2$ )
  - Amidogen (NHD)
  - Di-fluoromethyl ( $\text{CHF}_2$ )

# Possible precursors of Glycine



# Road to Glycine

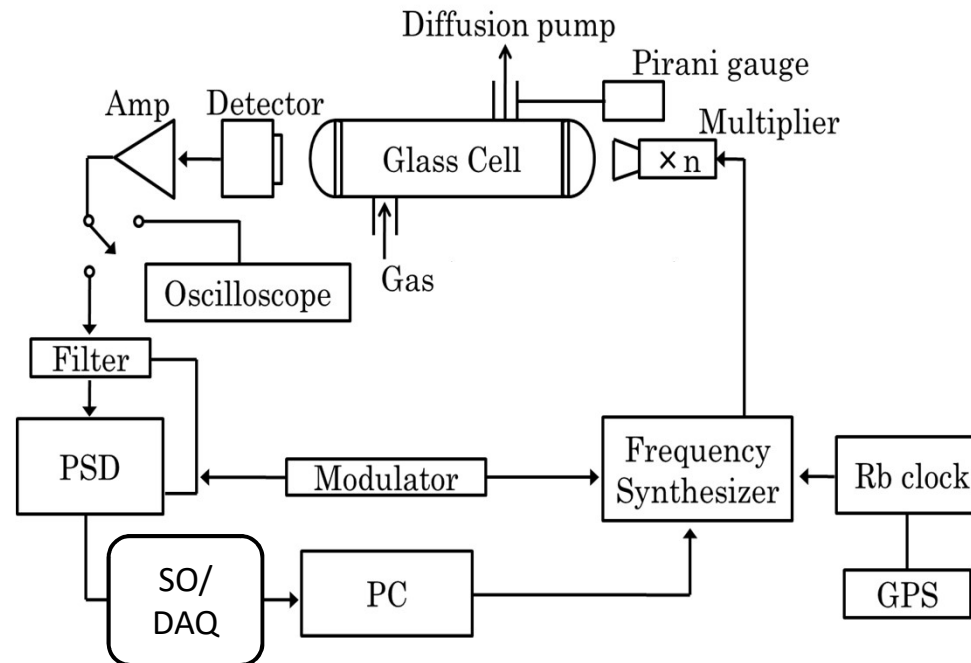


# Keys for spectroscopy of short-lived species

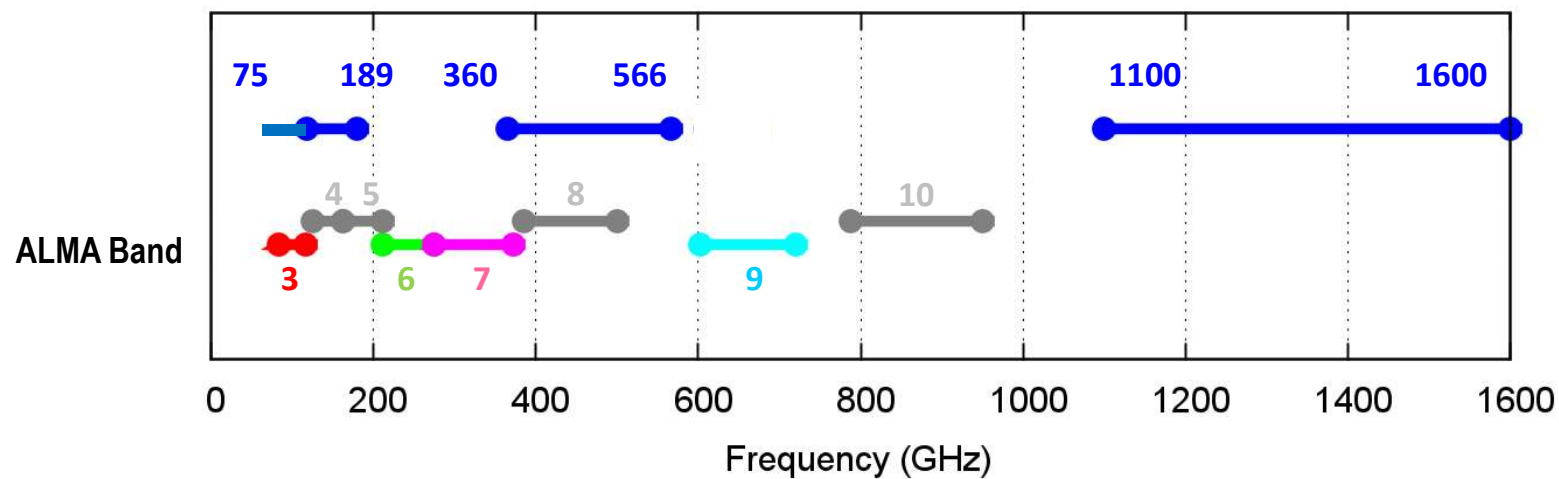
- Spectrometer sensitivity
- Production Efficiency
- Prediction Accuracy
- ...

# Spectrometer Sensitivity

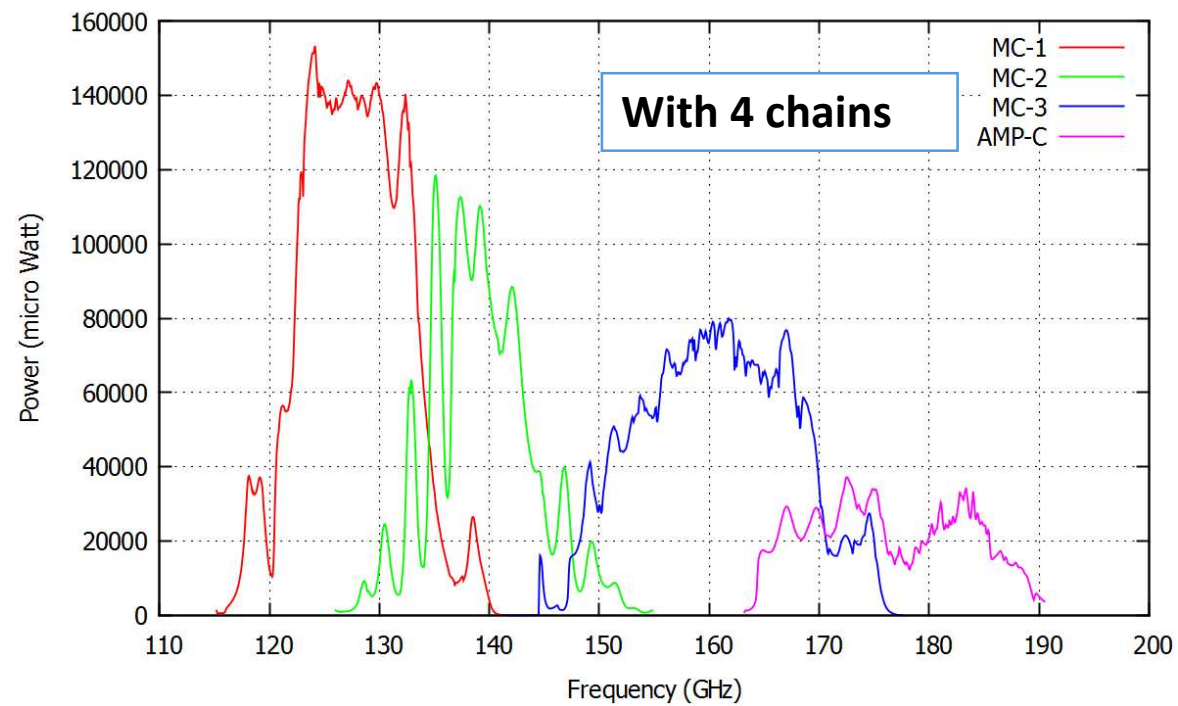
- Source noise vs Detector noise



# Frequency coverage

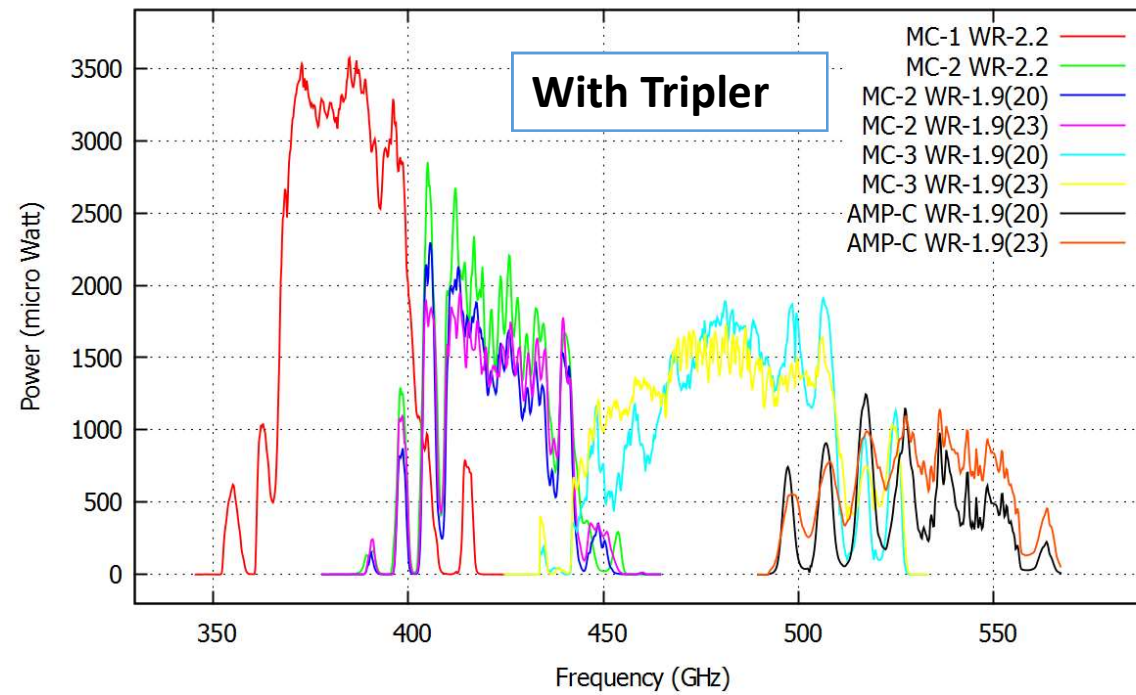


# Frequency coverage

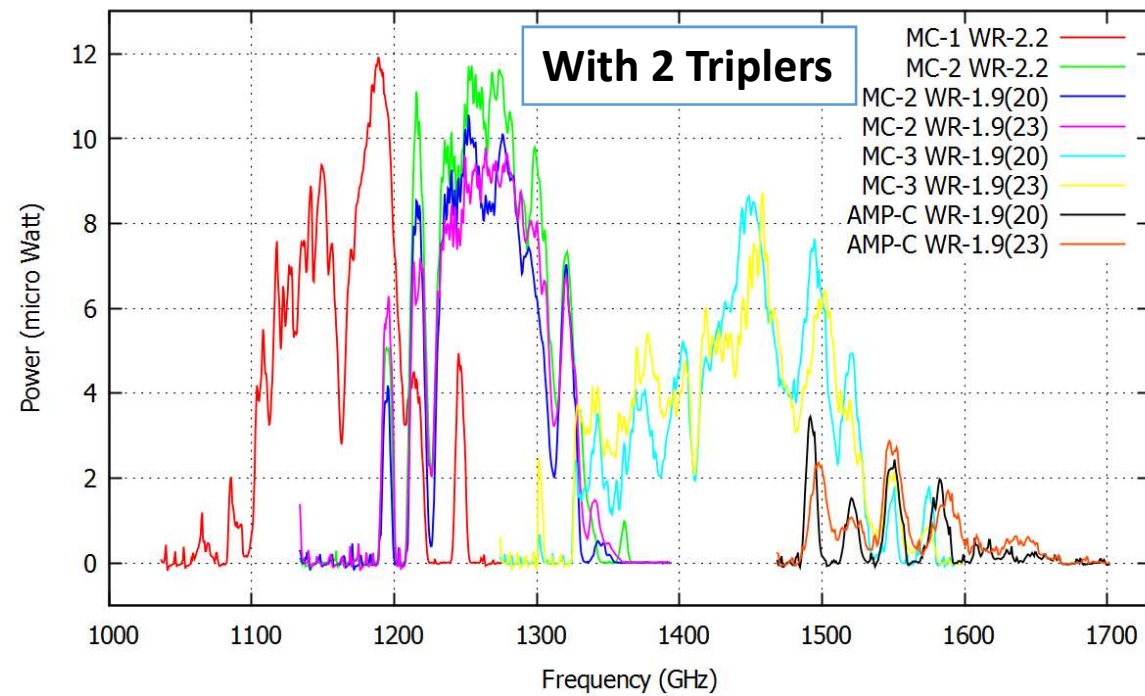




# Frequency coverage



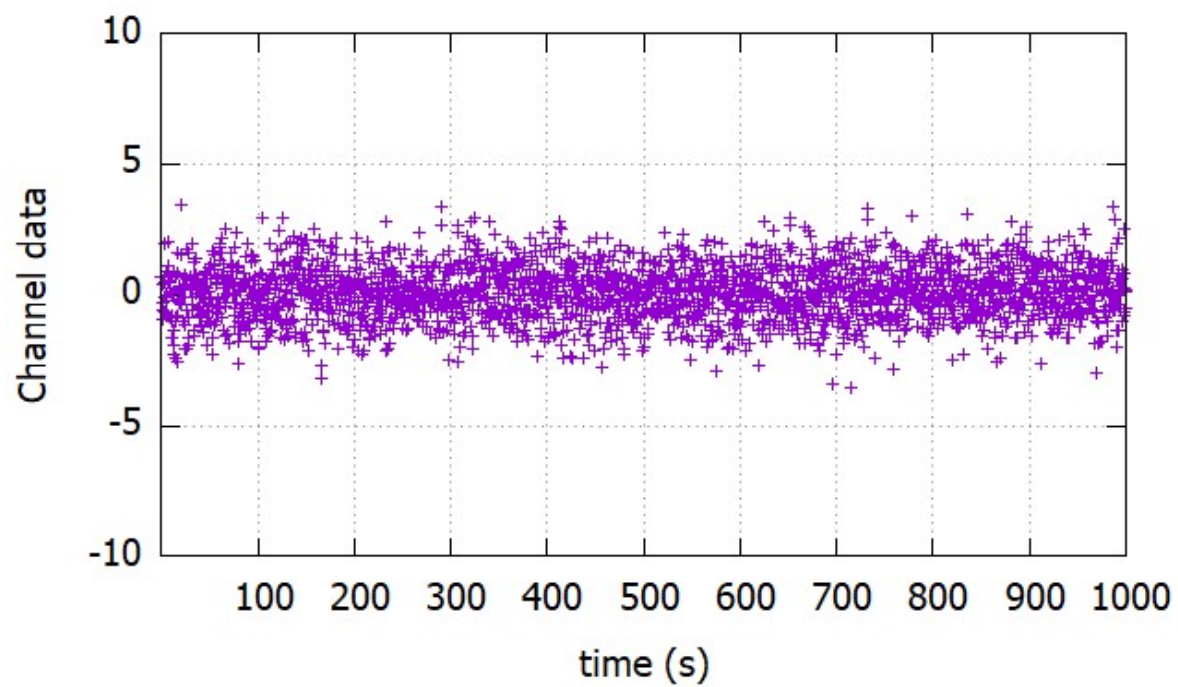
# Frequency coverage



# Other concerns

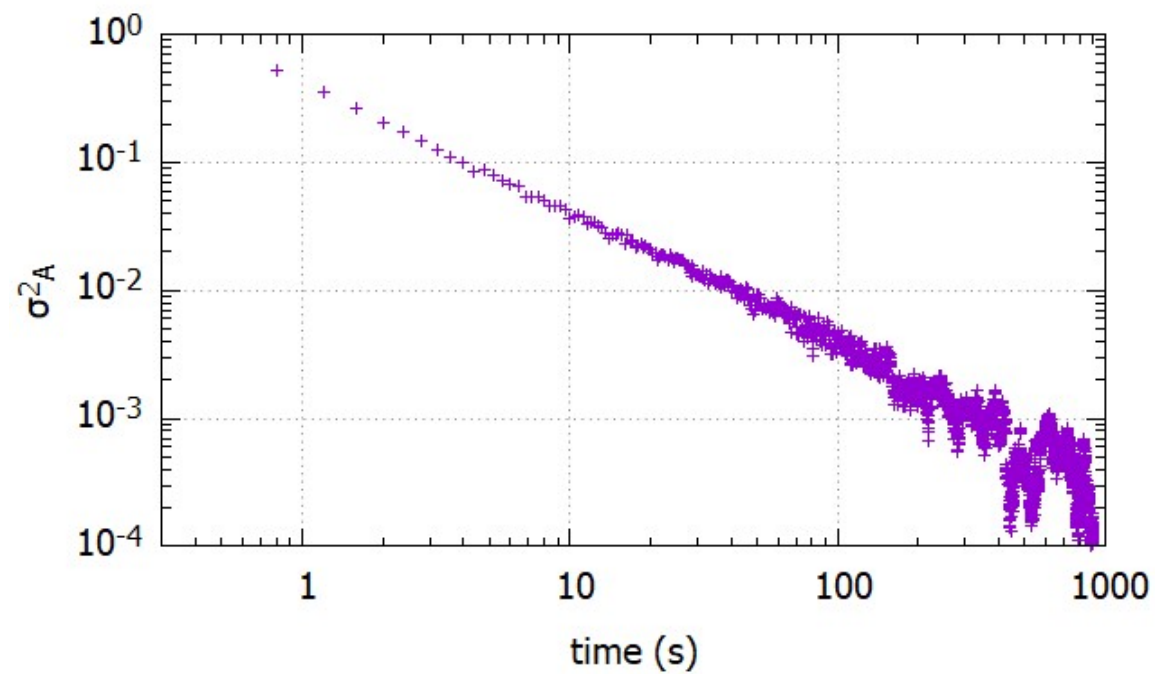
- Baseline distortion
- Discharge noise
- Earth magnetic field cancellation

# Channel data vs Time (Accm. #)

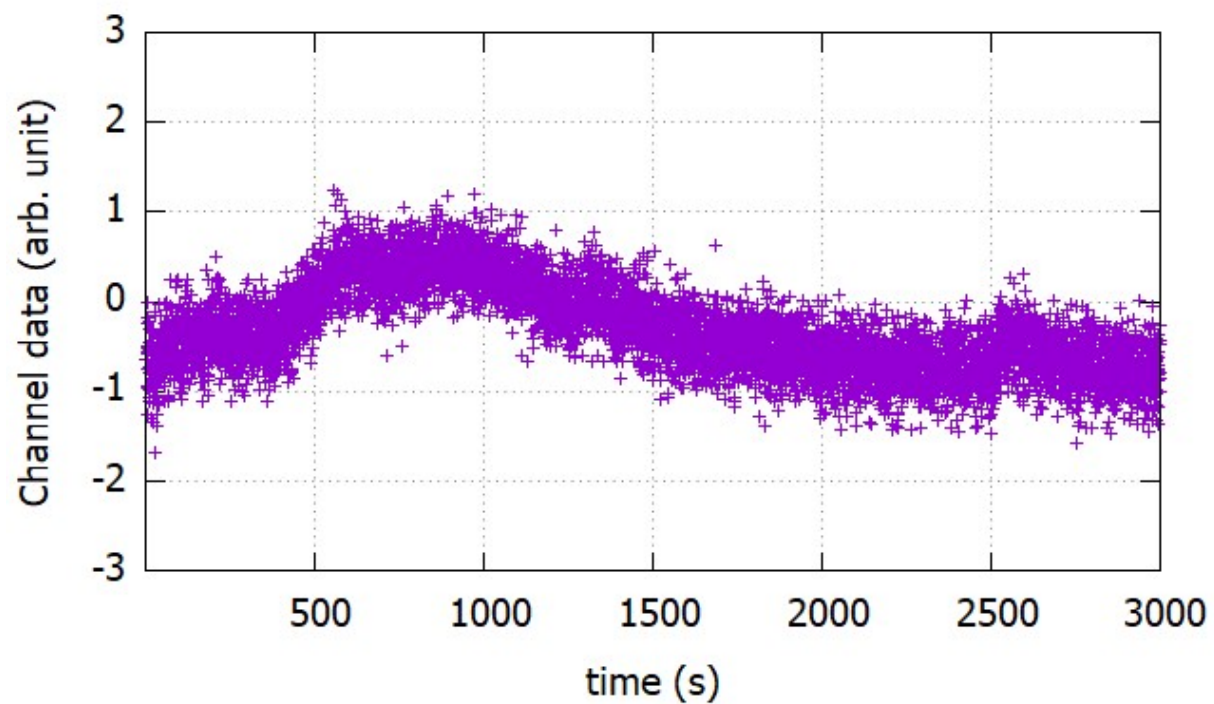


# “Spectroscopic Allan Variance”

after R. Schieder



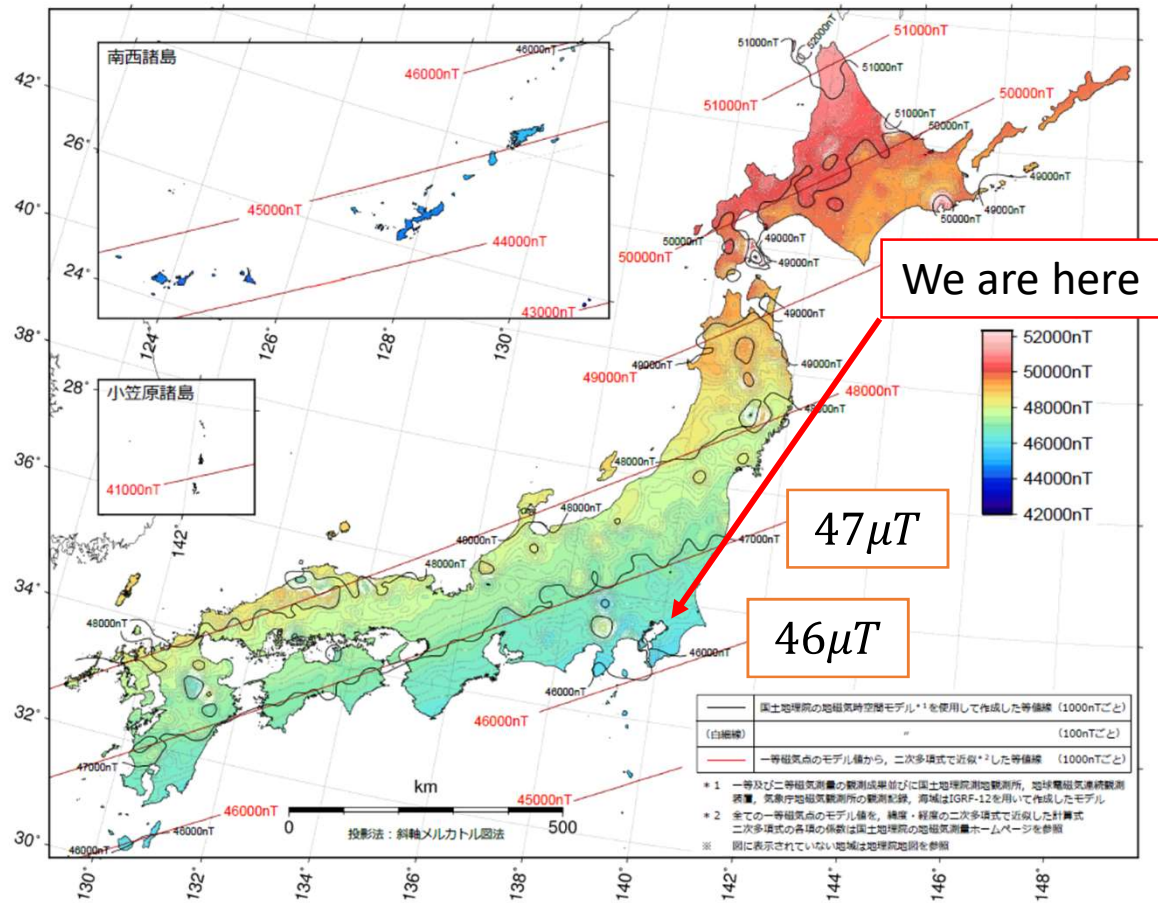
# Channel data vs Time (Accm. #)



# Other concerns

- Baseline distortion
- Discharge noise
- Earth magnetic field cancellation

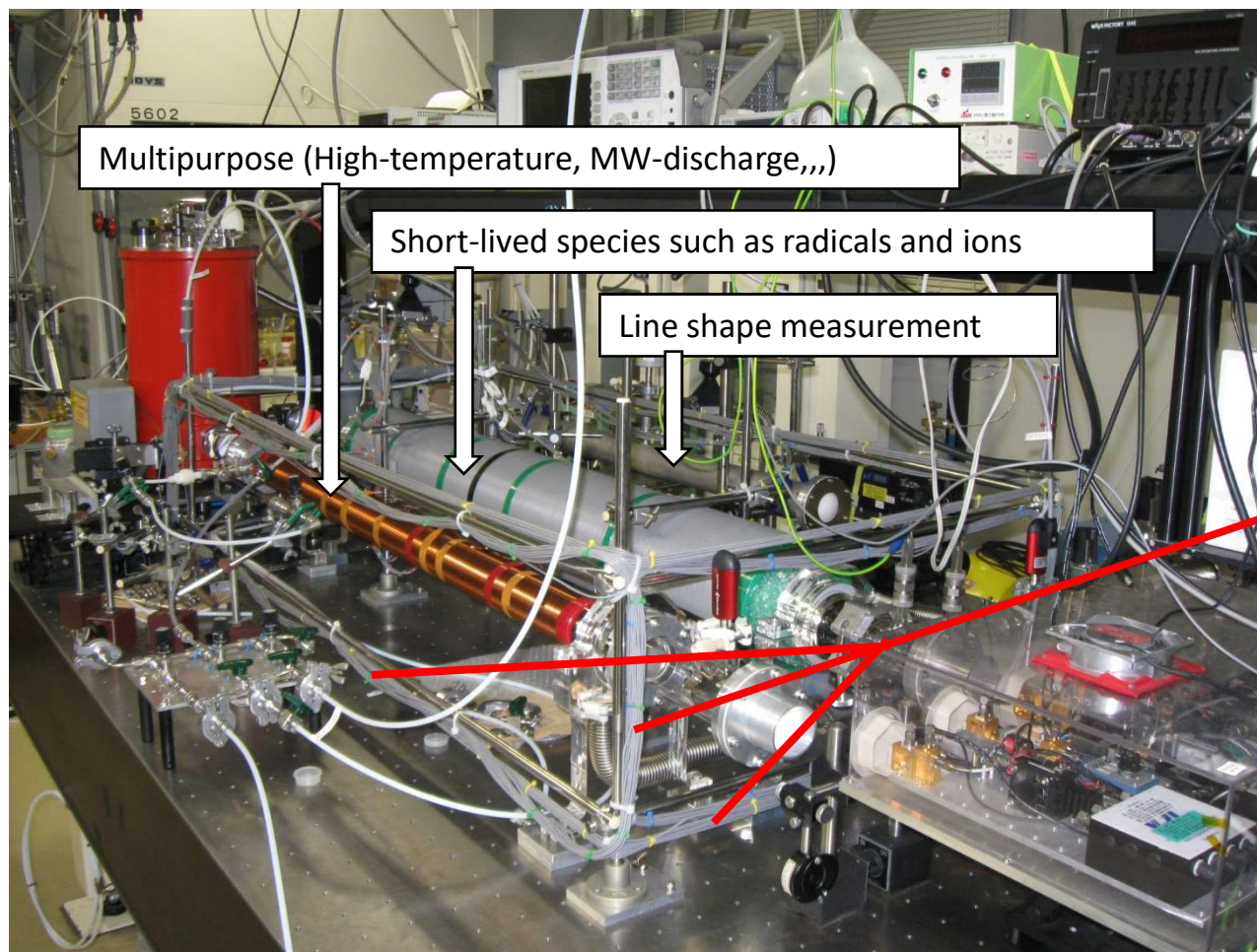
# Earth magnetic field



<http://www.gsi.go.jp/common/000148086.pdf>



# Spectrometer in Toho University (2018)



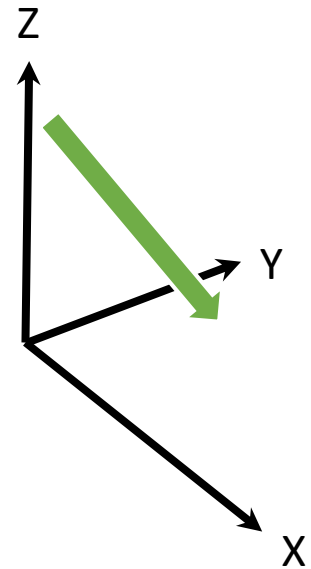
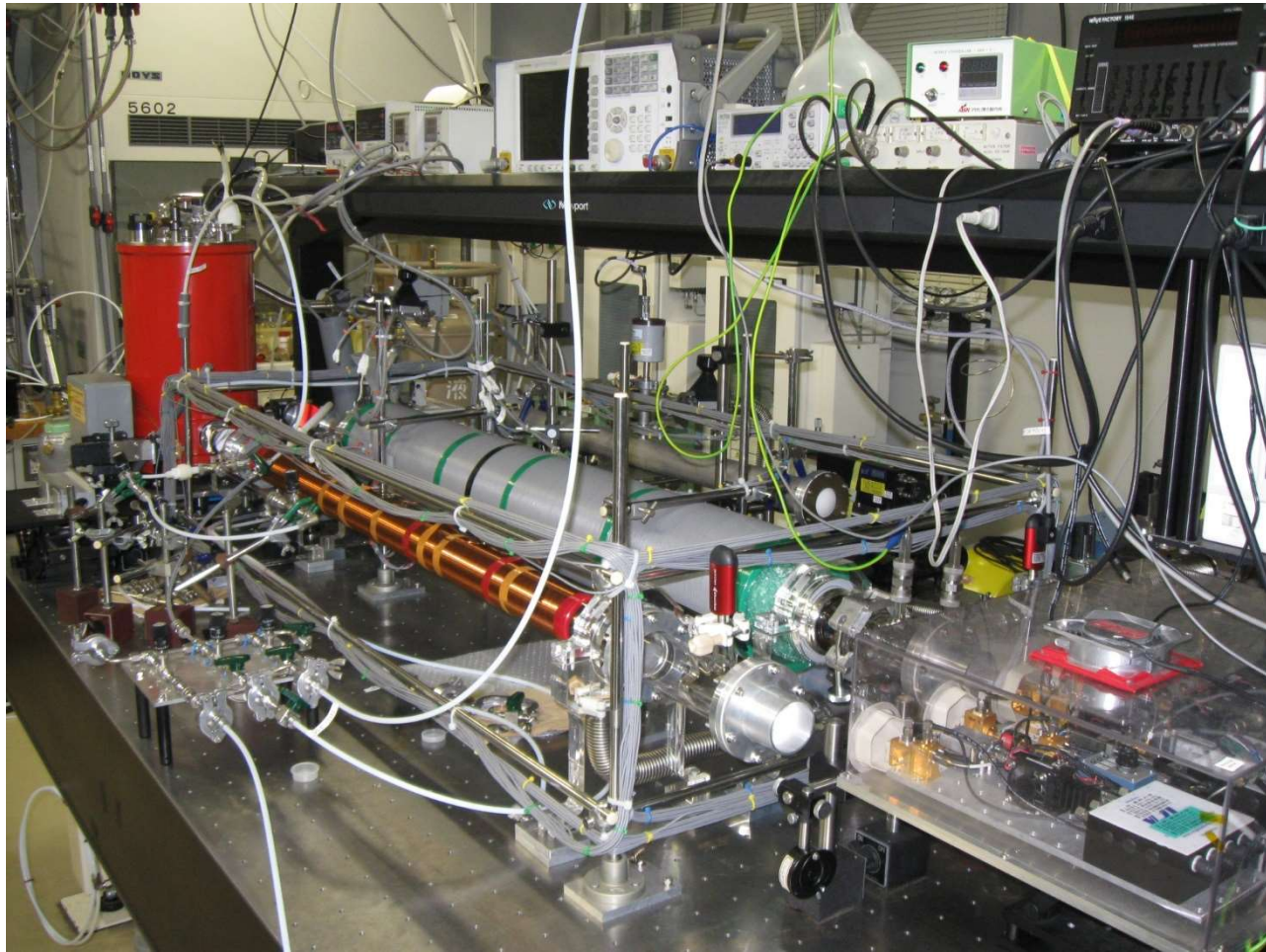
Multipurpose (High-temperature, MW-discharge,,,) ↓

Short-lived species such as radicals and ions ↓

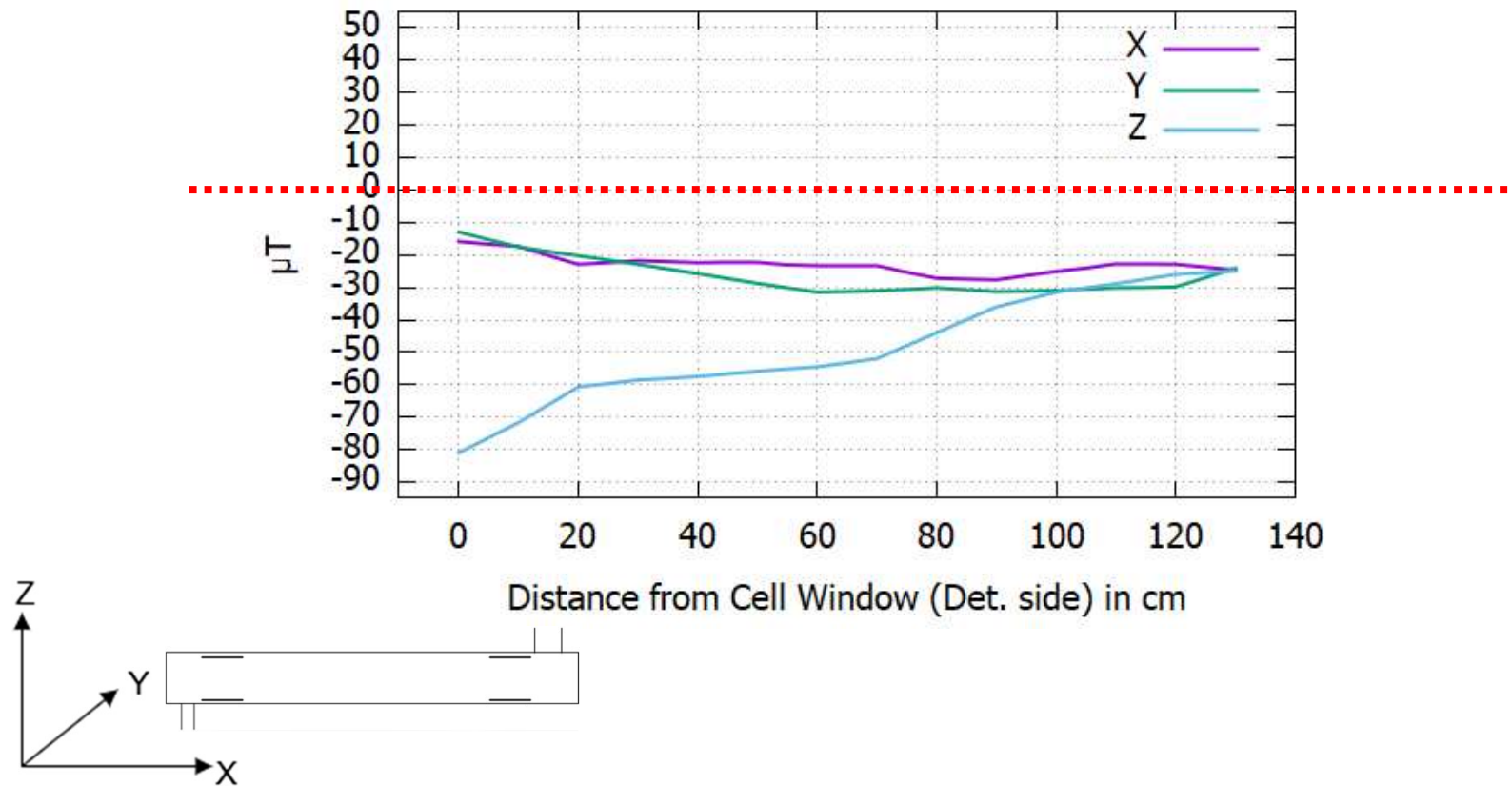
Line shape measurement ↓

**3 Pairs of  
Rectangular coils**

# Spectrometer in Toho University (2018)



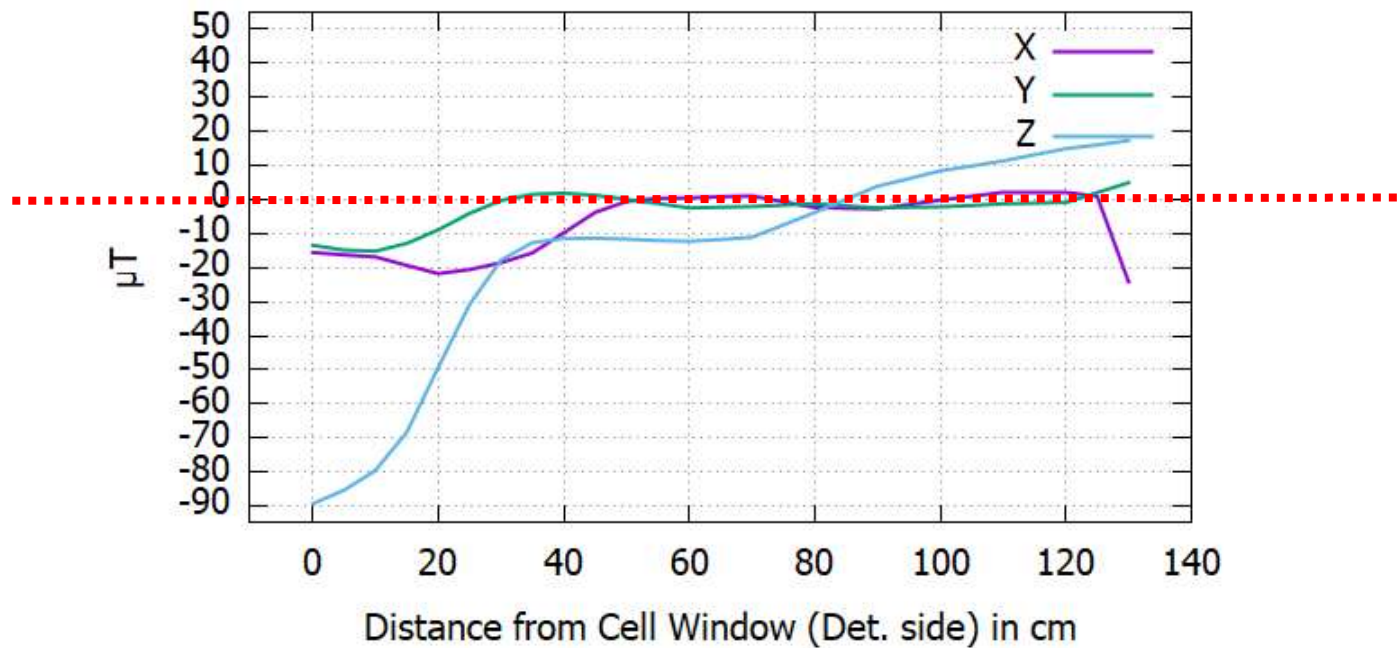
# Magnetic field around spectrometer





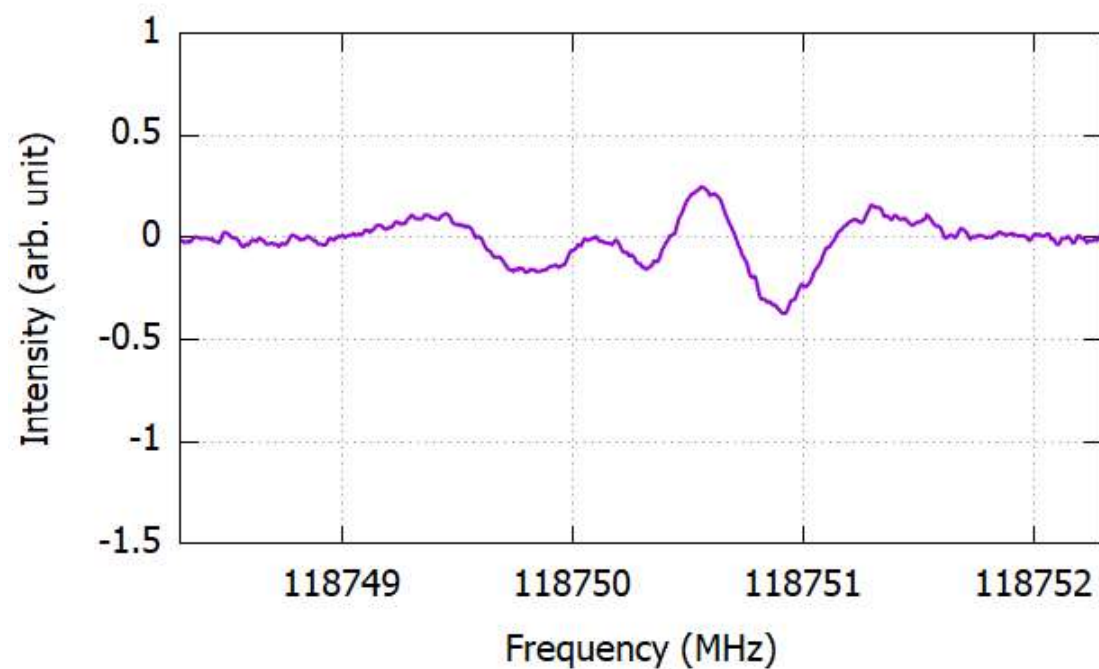
# Magnetic field around spectrometer

After optimizing current setting of coils



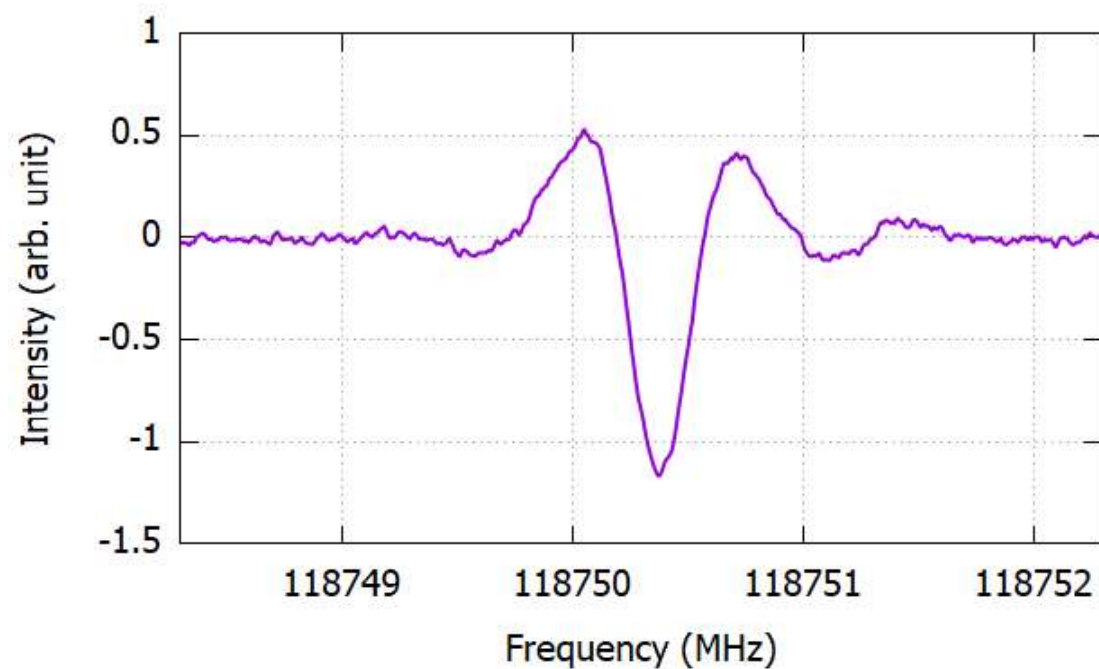
# Oxygen line @118 GHz ( $1_1-1_0$ )

Before optimizing current setting of coils



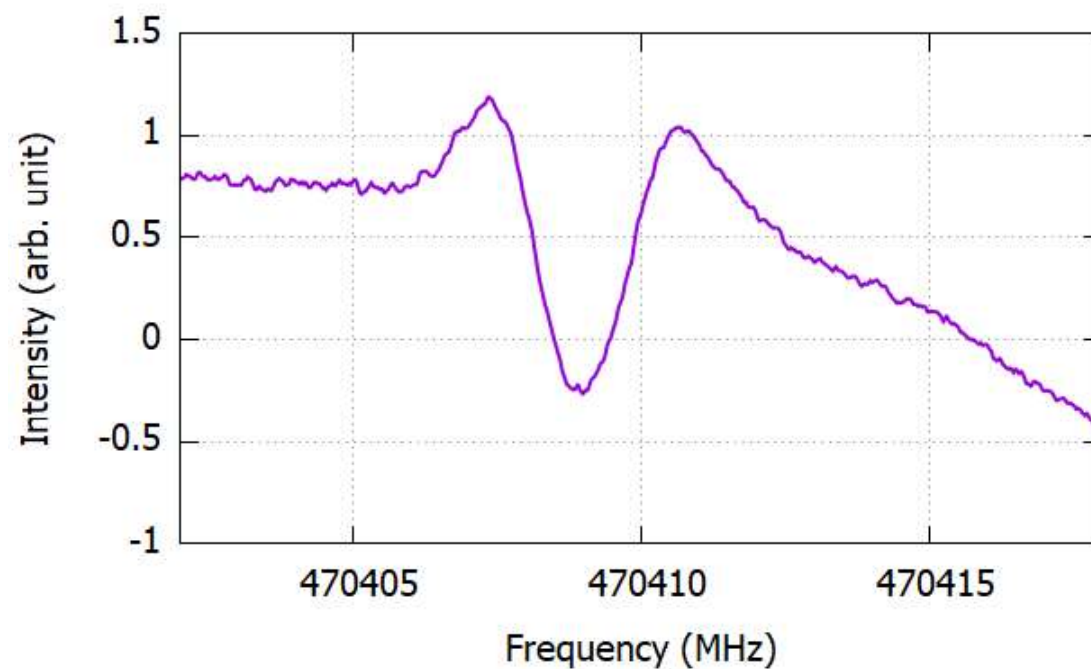
# Oxygen line @118 GHz ( $1_1-1_0$ )

After optimizing current setting of coils



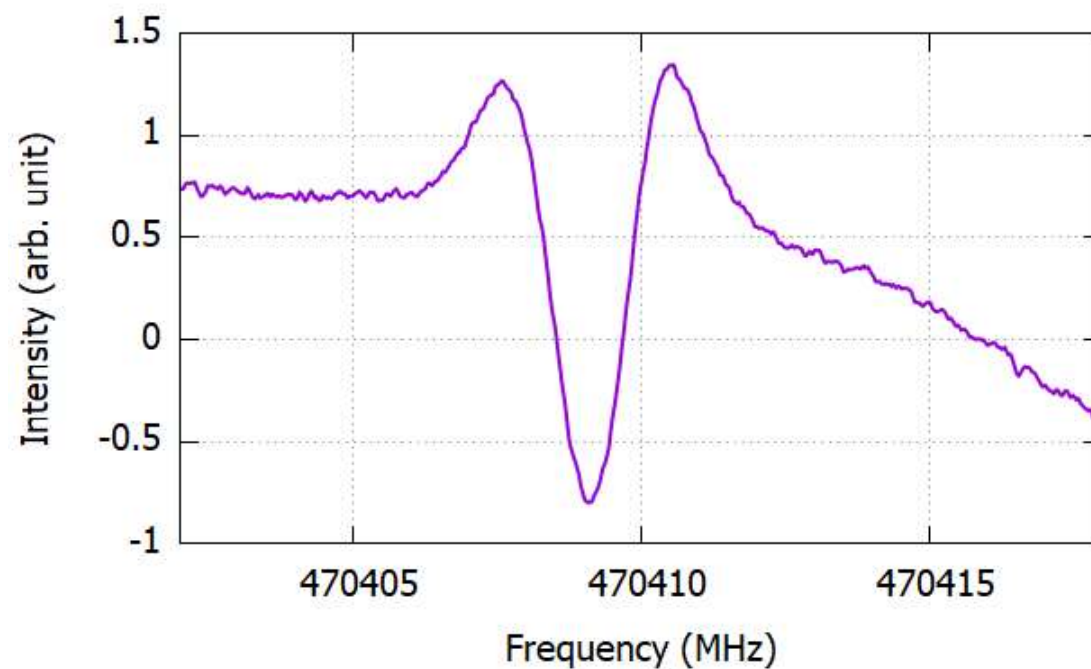
# NH<sub>2</sub> radical @470 GHz ( $1_{10}-1_{01}$ )

With current setting for observing oxygen line @ 118 GHz

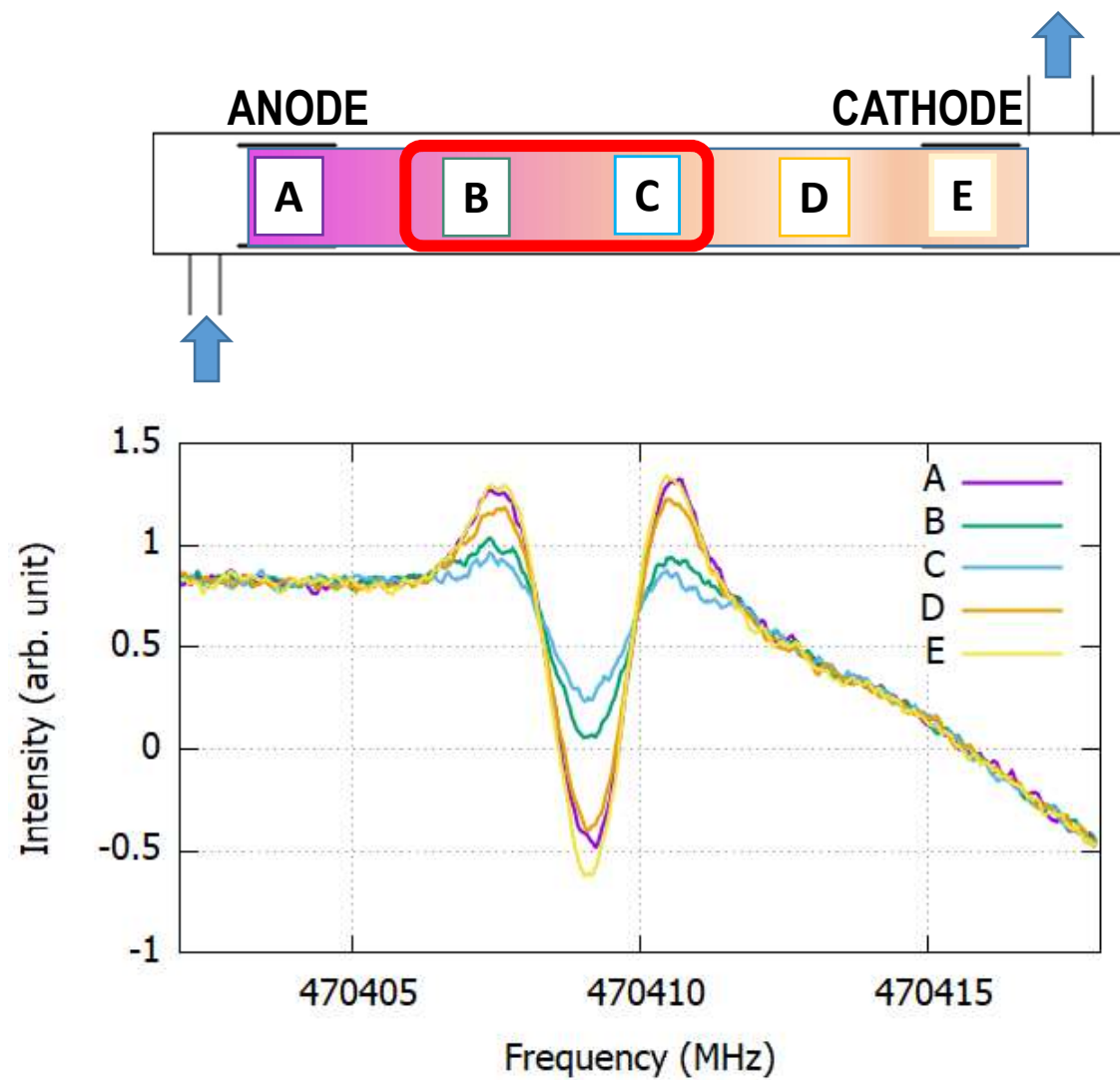


# NH<sub>2</sub> radical @470 GHz ( $1_{10}-1_{01}$ )

After re-optimization current setting of coils

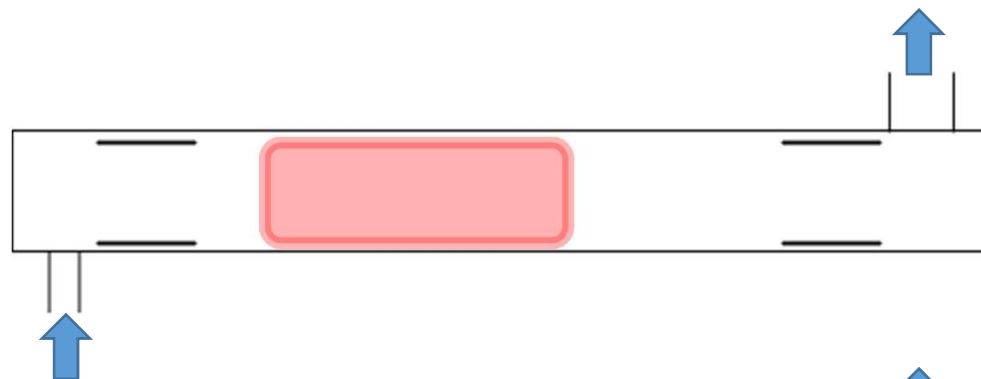




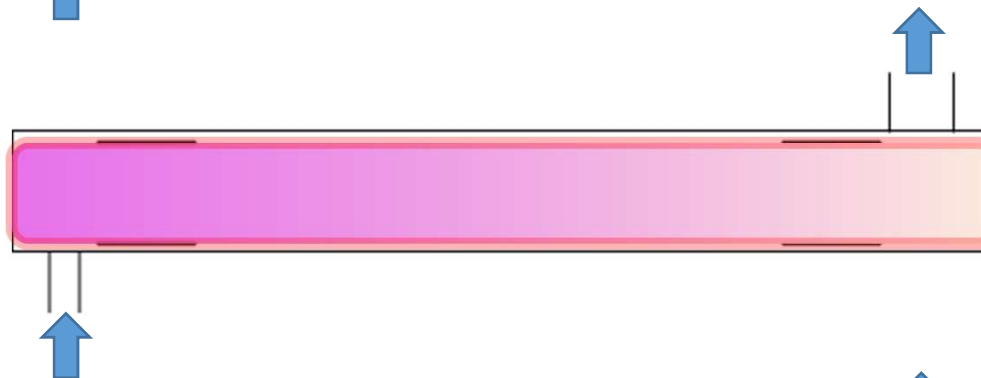


NH<sub>2</sub> radical @470 GHz ( $1_{10}-1_{01}$ )

$\text{NH}_2$ ,  $\text{CD}_2$   
DC-glow disch.



$\text{O}_2$   
Stable species



$\text{CHF}_2$   
Rxn. at outside



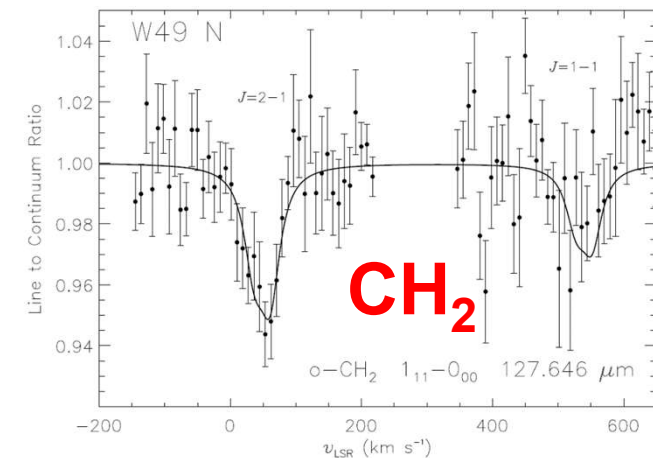
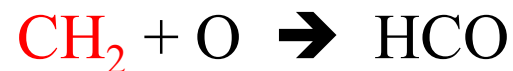
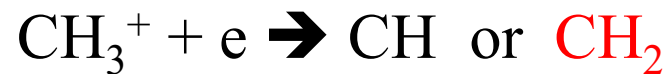
# $\text{CD}_2$ : (Deuterated) Methylene radical

Difficult to produce enough amount for spectroscopy

# Methylene radical in space

- Relatively abundant in diffuse/dense interstellar cloud
- Most of astronomically important transitions lie in **terahertz region**

- Major production/Destruction Process

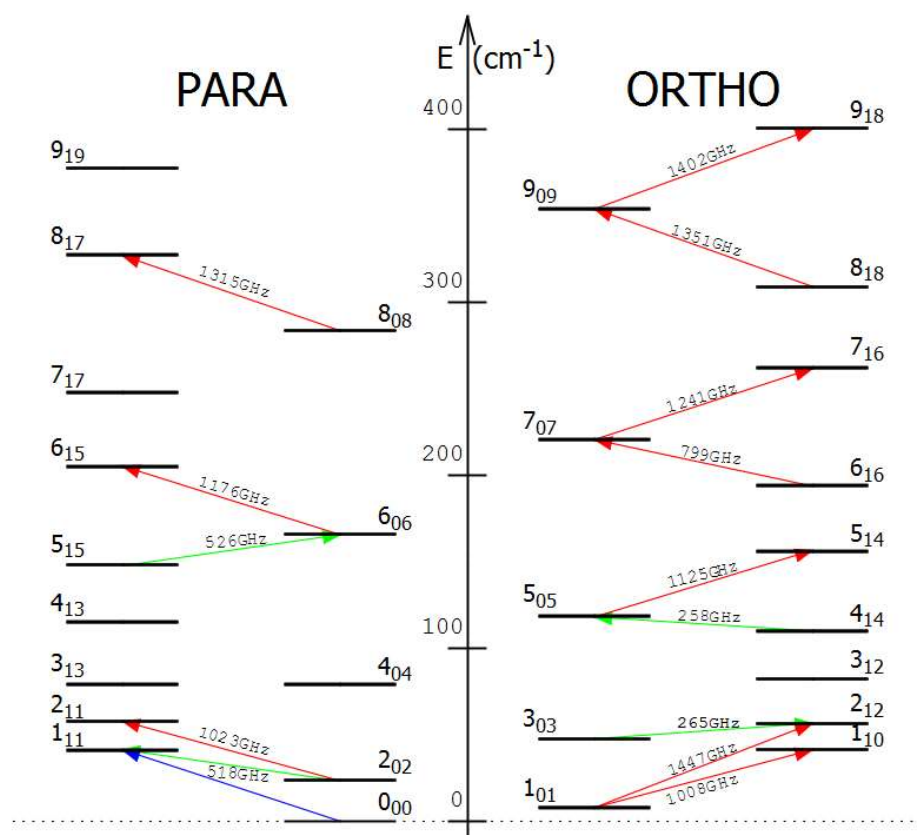


$1_{11}-0_{00}$  transition ( 2350 GHz)  
E. T. Polehampton et al. (2013)

# Methylene radical in the ground state

CH <sub>2</sub> (THz Spectroscopy)	S. Brunken et al. (2004).
(Global fit)	S. Brunken et al. (2005).
CHD (LMR)	J. Nolte et al. (1994).
(THz Spectroscopy)	H. Ozeki et al. (2011).
CD <sub>2</sub> (LMR)	K. M. Evenson et al. (1984).
(MW Spectroscopy)	H. Ozeki et al (1996).
<b>(THz Spectroscopy)</b>	<b>Present Work</b>

# Energy level diagram of CD<sub>2</sub>

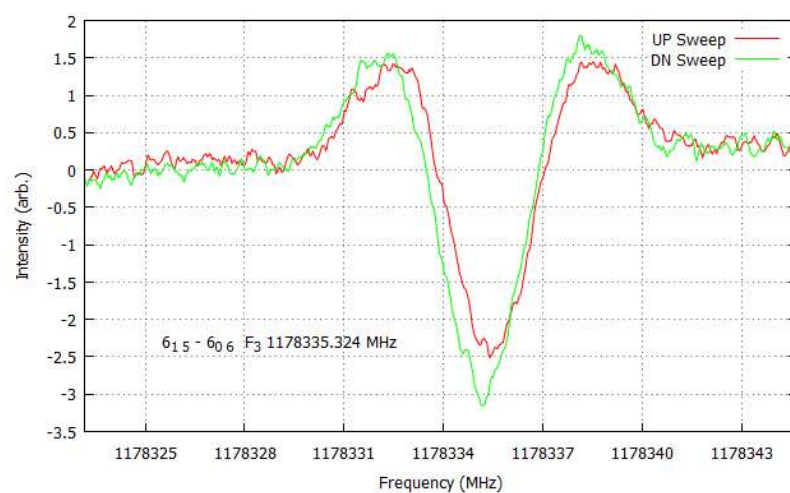


# of	Transitions	Lines
Microwave	4	19
Terahertz	10	29

→ Terahertz  
→ Microwave  
→ Astronomically Important Transitions

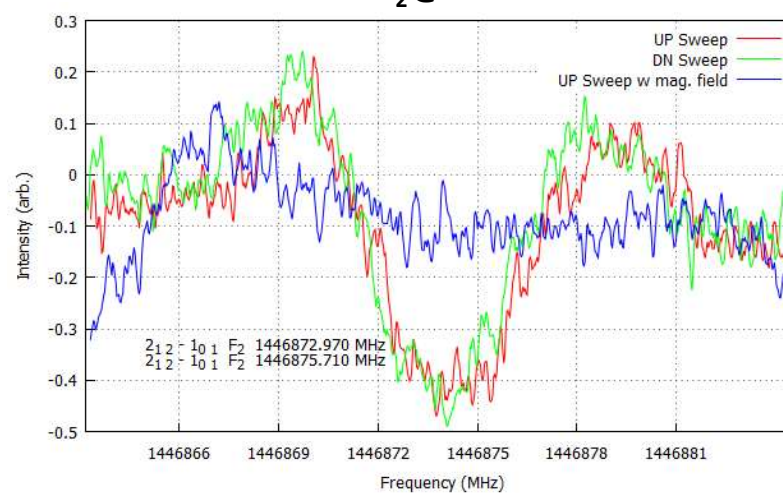
# CD<sub>2</sub> : Transitions with $K_a = 1 - 0$

Para-CD<sub>2</sub> @ 1178 GHz



$6_{15} - 6_{06} F_3$

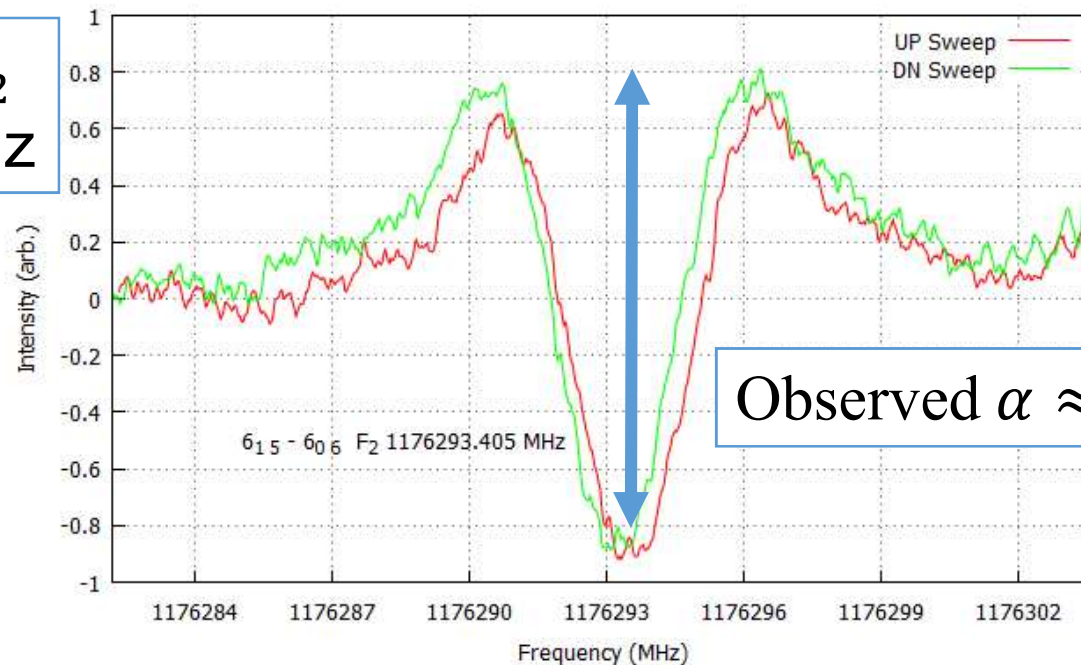
Ortho-CD<sub>2</sub> @ 1446 GHz



$2_{12} - 1_{01} F_2$

# Estimation of production efficiency

$6_{15} - 6_{06} F_2$   
@1176 GHz



Observed  $\alpha \approx 10^{-5} \text{ cm}^{-1}$



# Estimation of production efficiency

- Experimental condition

- Total Pressure 5 Pa
- Cell length 150 cm

- Observed optical thickness  
➔  $\sim 1.5 \times 10^{-3}$

- Calculated spectral Intensity

Log (I) @ 300 K (nm <sup>2</sup> MHz)	$F'' - F'$
-2.409	7-7
-2.547	5-5
-2.483	6-6
-2.00	(total)

- Expected optical thickness @  
100 % production efficiency  
➔  $\sim 6.7$

$$\eta \sim 0.02 \%$$

# Demand for more efficient production

- $\text{CD}_4$  in Ar  $\xrightarrow{\text{DC-glow disch.}}$   $\text{CD}_2$
- $\text{CD}_4 + \text{F} \longrightarrow \text{CD}_2$
- $\text{CD}_4 \xrightarrow{\text{Ar}^* \text{ or He}^*} \text{CD}_2$
- $(\text{CD}_2\text{CO})_2 \xrightarrow{\Delta (800\text{K})} \text{CD}_2\text{CO} \xrightarrow{\text{DC-glow disch.}} \text{CD}_2$
- $(\text{CD}_3\text{CO})_2\text{O} \xrightarrow{\Delta (800\text{K})} \text{CD}_2\text{CO} + \text{CD}_3\text{COOD} \xrightarrow{\text{DC-glow disch.}} \text{CD}_2$
- ???

# Molecular constants of CD<sub>2</sub>

Constants	Present Work	FIR-LMR <sup>a</sup>
A	1132826.442(110) <sup>b</sup>	1132820.6(18)
B	126805.268(34)	126804.1(34)
C	110760.770(24)	110760.1(34)
$\Delta_J$	2.80044(124)	2.771(19)
$\Delta_{JK}$	-141.1343(101)	-149.16(24)
$\Delta_K$	16795.21 <sup>c</sup>	16795.21(48)
$\delta_J$	0.65606(32)	0.669(7)
$\delta_K$	83.4 <sup>c</sup>	83.4(15)
$\Phi_J$	-0.0002631(43)	
$\Phi_{JK}$	-0.07739(47)	-0.057(5)
$\Phi_K$	588. <sup>c</sup>	588. <sup>c</sup>

# Molecular constants of CD<sub>2</sub> (Cont'd)

Constants	Present Work	FIR-LMR <sup>a</sup>
$\varepsilon_{aa}$	10.674(240)	8.5(11)
$\varepsilon_{bb}$	-76.190(48)	-78.3(11)
$\varepsilon_{cc}$	-62.425(65)	-62.4(119)
$\alpha$	7770.115(147)	7759.3(9)
$\alpha_K$	-40.827(168)	
$\beta$	1222.897(149)	1216.6(9)
$a_F(D)$	-2.903(141)	
$T_{aa}(D)$	5.036(57)	
$T_{bb}(D)-T_{cc}(D)$	-0.076	

<sup>a</sup> K. M. Evenson et al (1984).

<sup>b</sup> 1 $\sigma$

<sup>c</sup> fixed

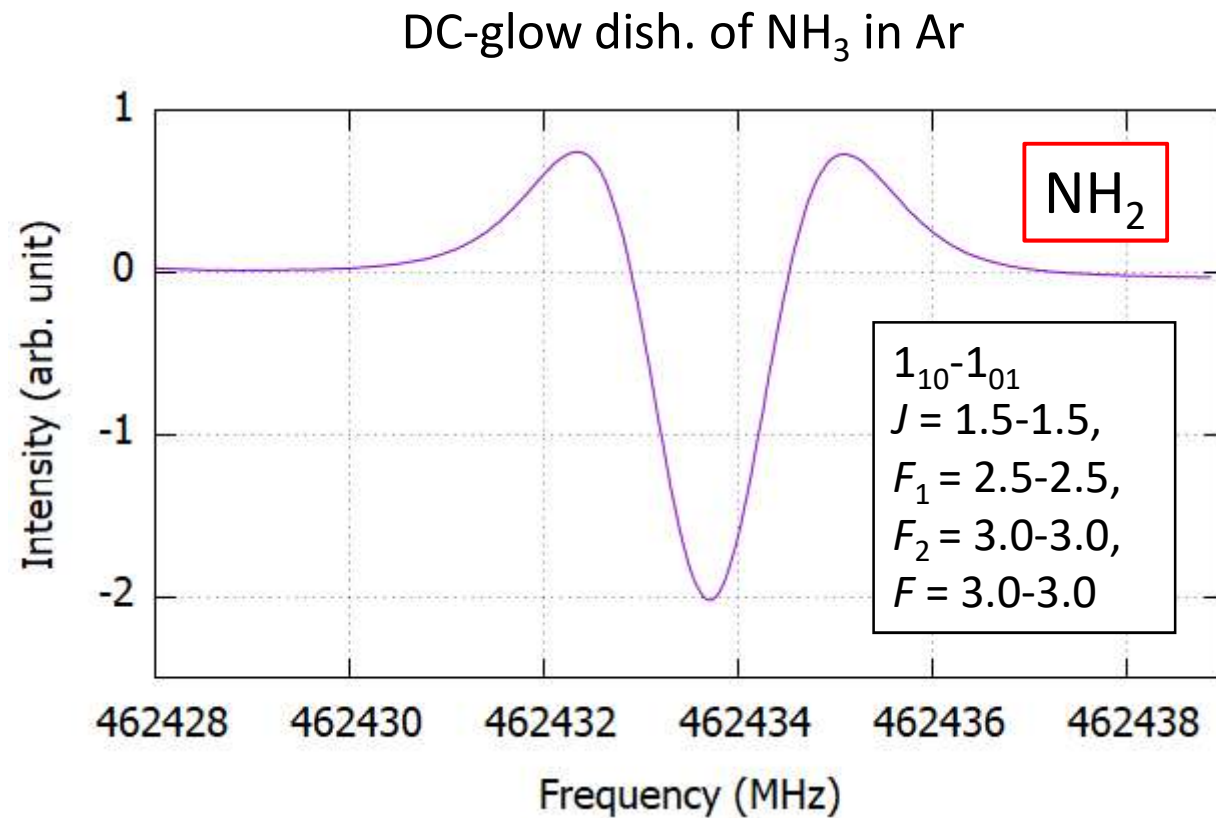
NHD :  
(Mono-Deuterated)  
Amidogen radical

Relatively easy to produce, but hard to analyze the spectra

# Mono-deuterated Amidogen NHD

- Amidogen : Relatively easy to produce  
: Light molecule  
→ Major transitions lie in the terahertz region  
  
NHD : 4 spin system :  $S + I(N) + I(H) + I(D)$   
→ Gives complex fine & hyperfine spectra

Amidogen : easier to produce than Methylene



Production efficiency  $\eta > 1\%$

# Complex fine & hyperfine spectra

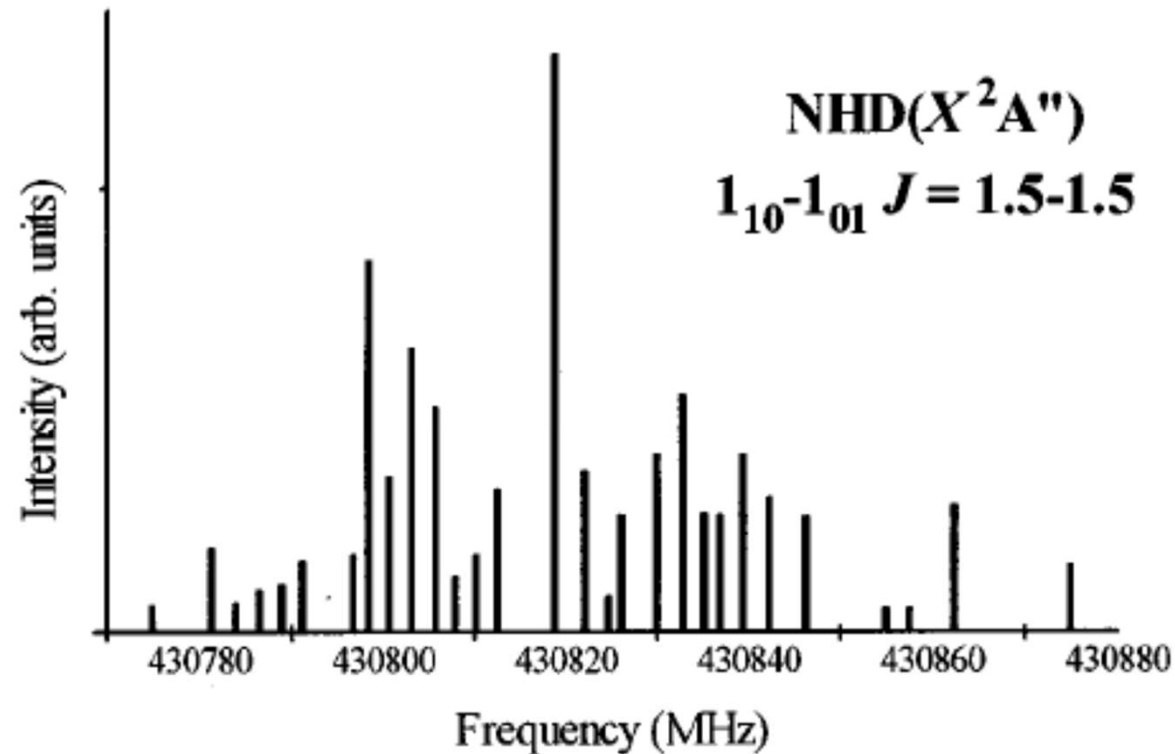
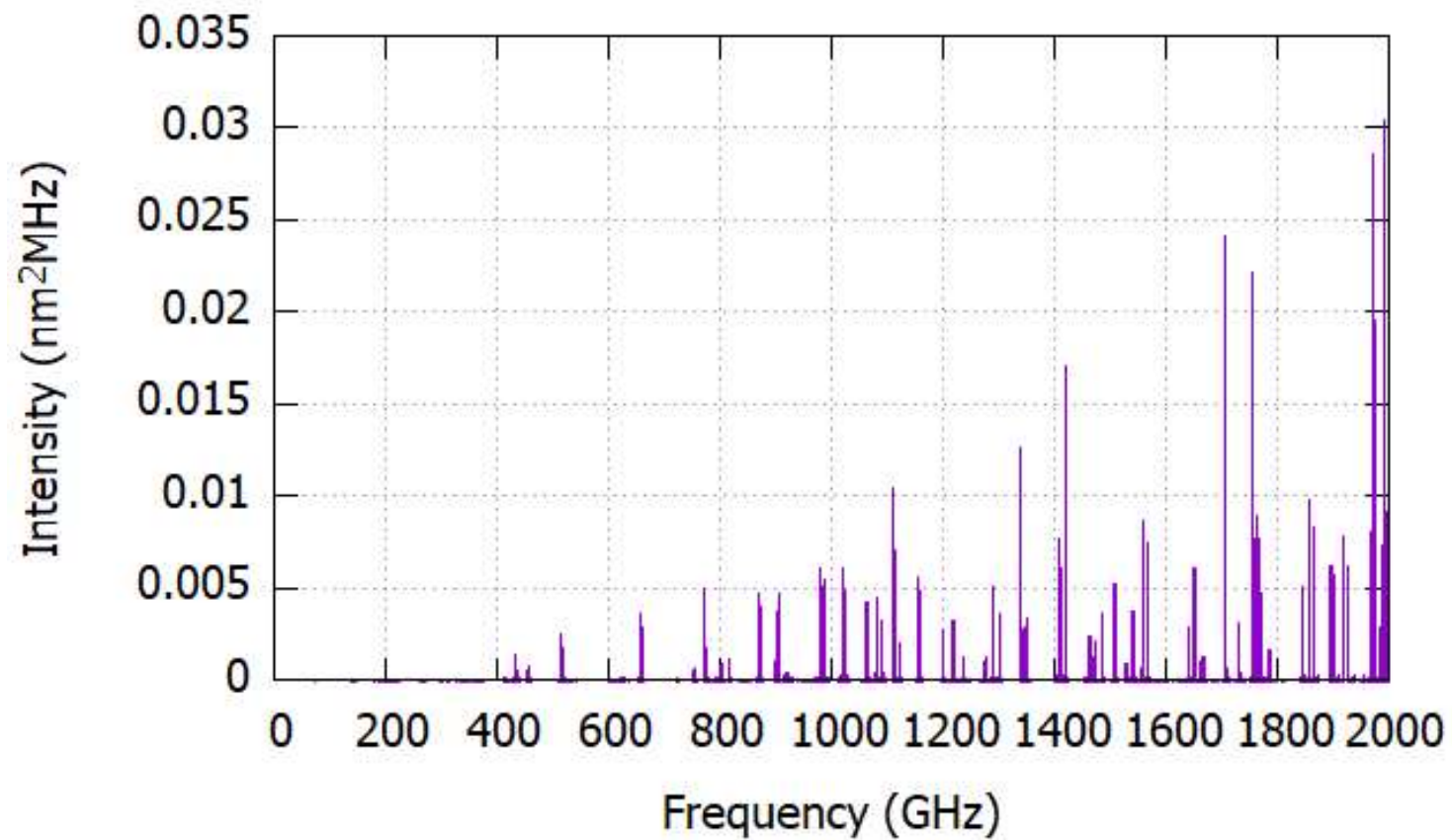


FIG. 2. The stick diagram of  $1_{10}-1_{01}$   $J=1.5-1.5$  transition of NHD ( $X^2A''$ ) in the 430 GHz region with observed relative intensities.

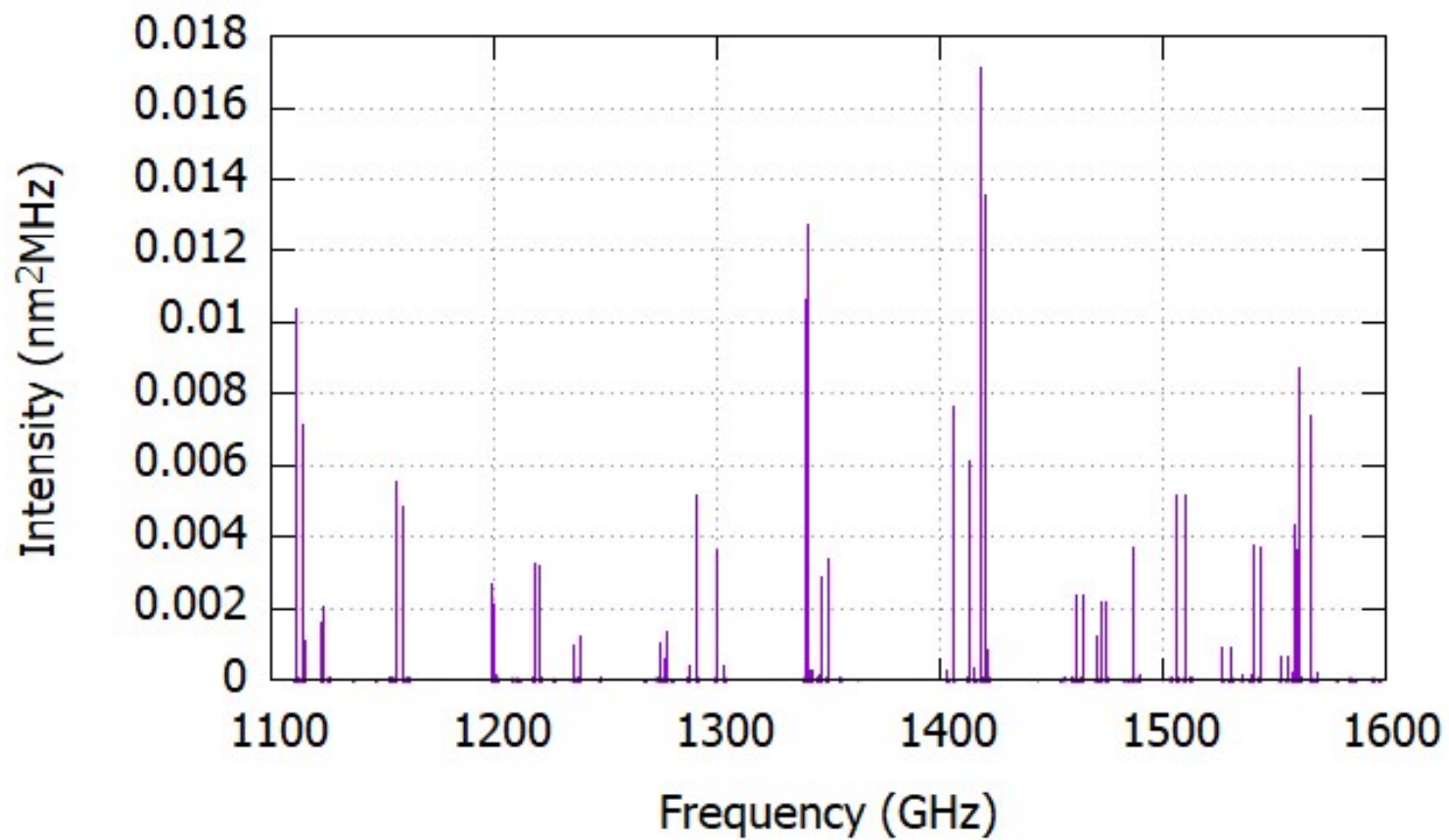
K. Kobayashi et al. (1997)



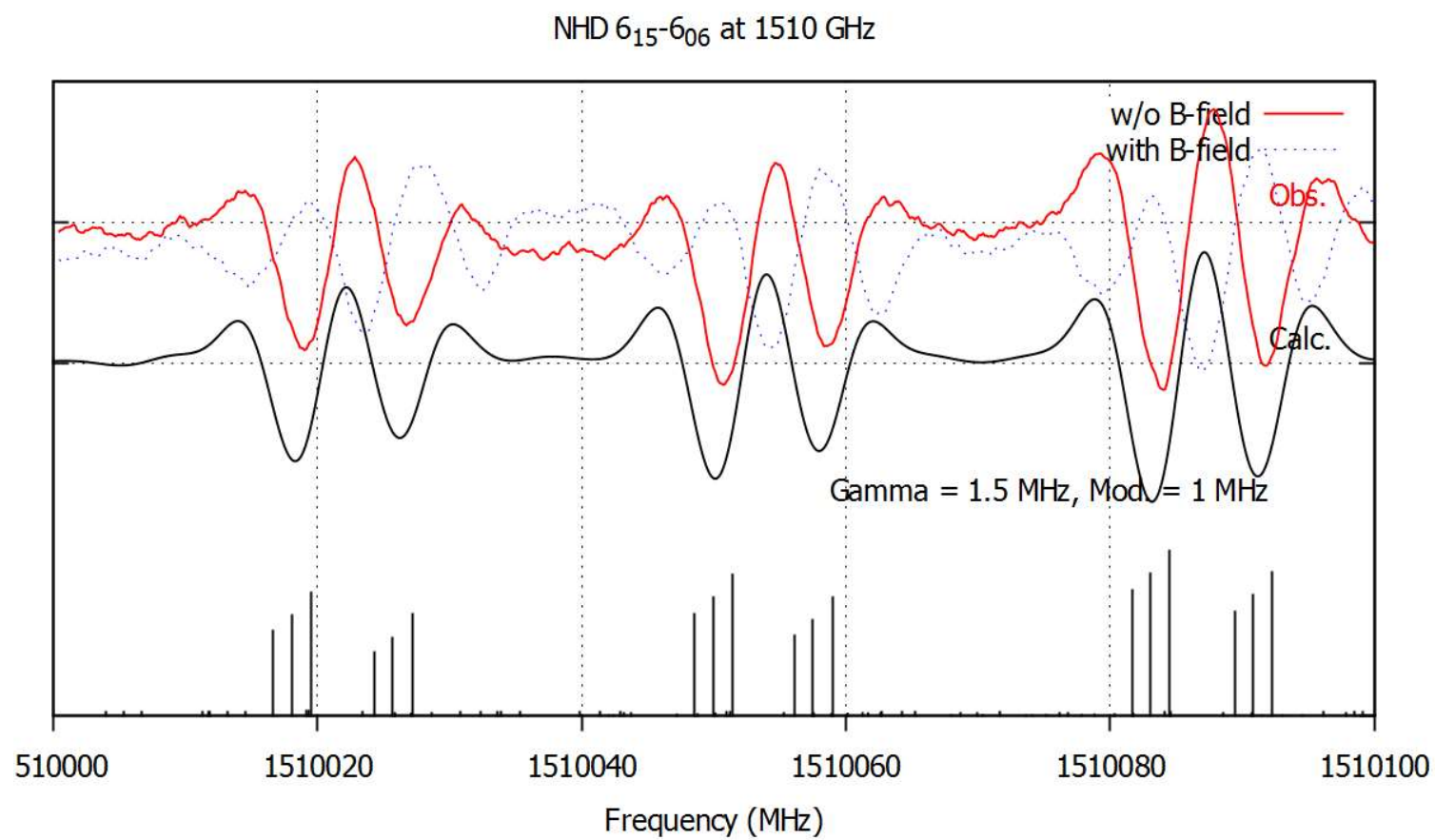
# NHD @ 300 K

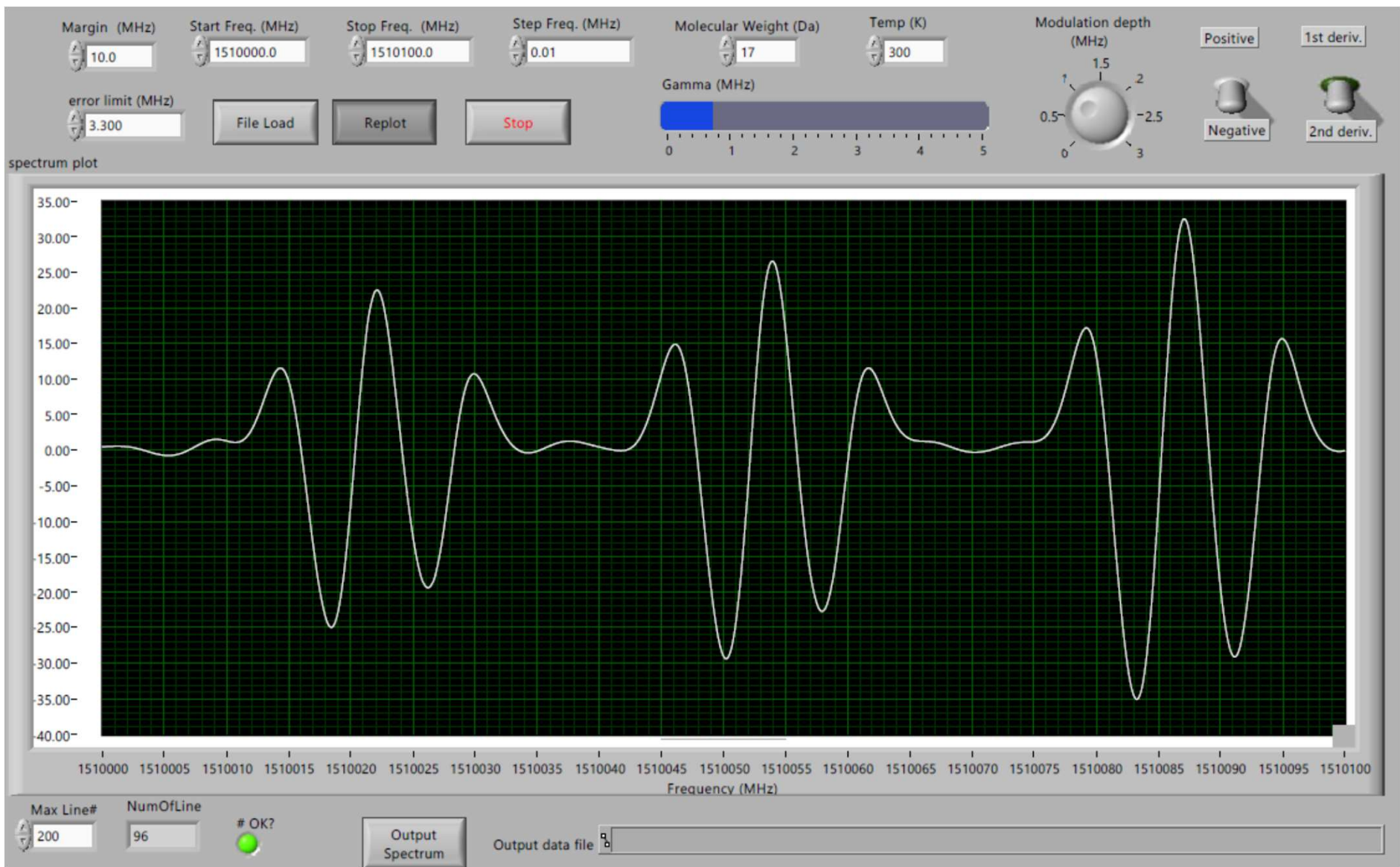


# NHD @ 300 K

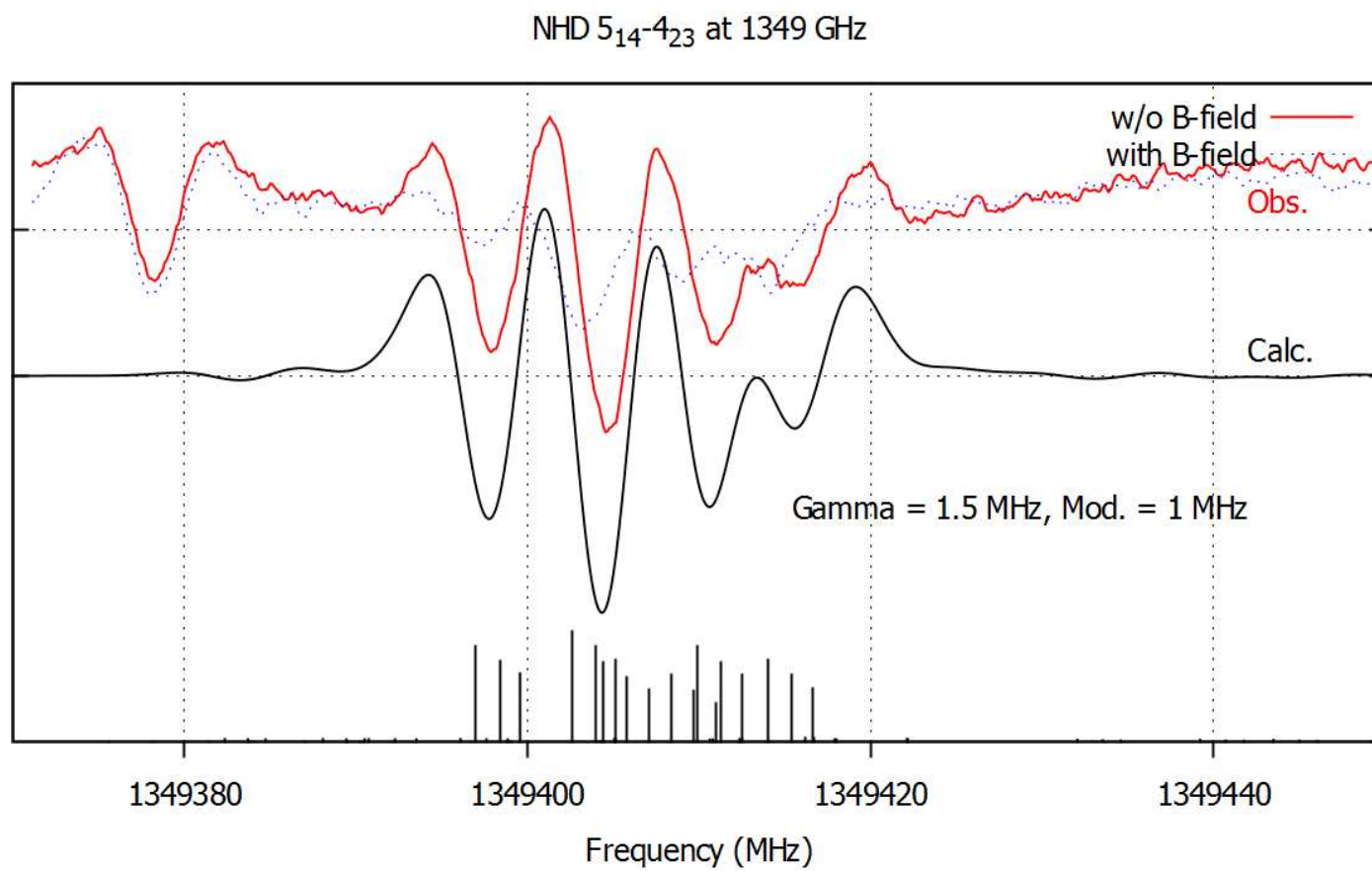


# Case 1 (OK)



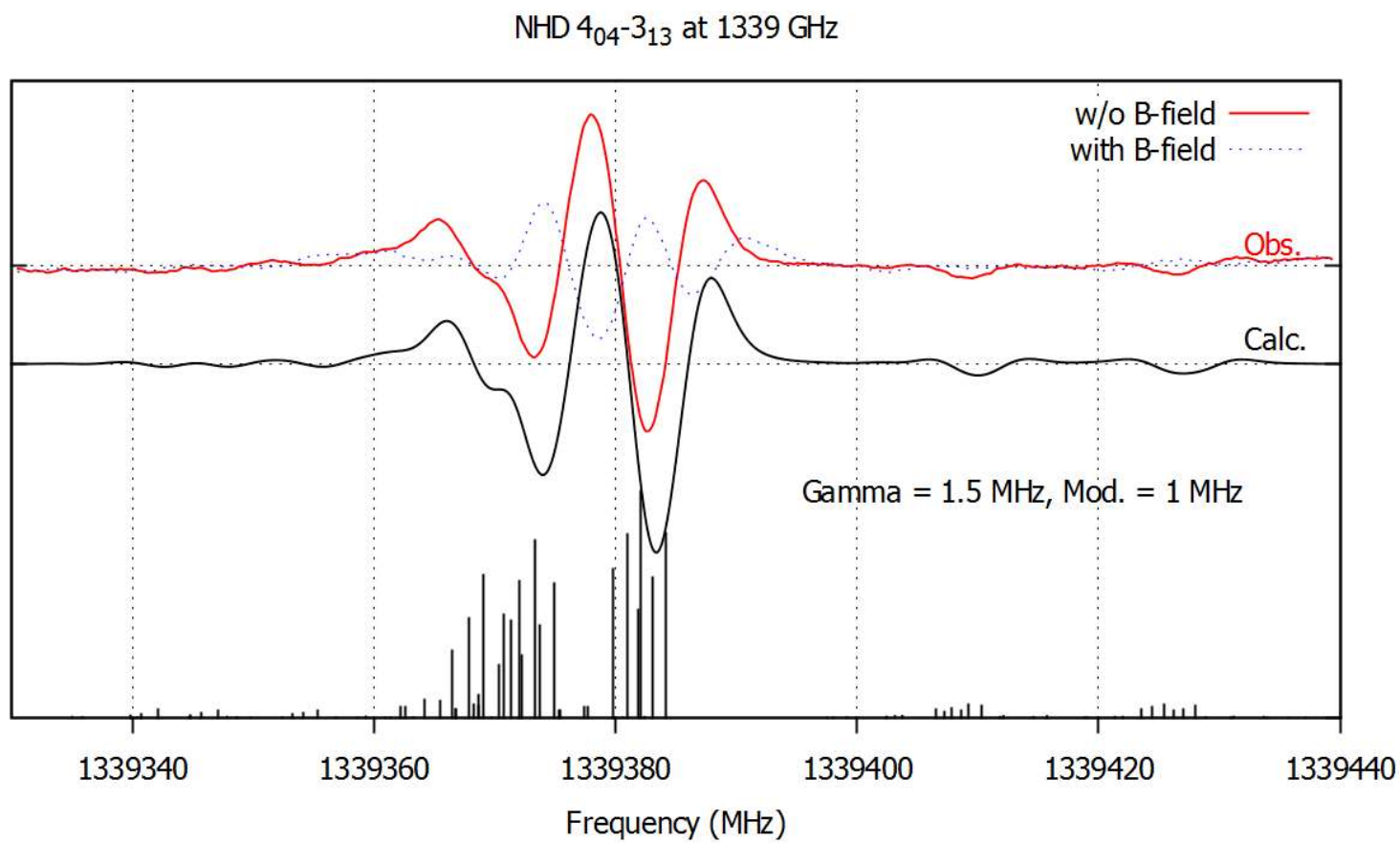


## Case 2 (NG)

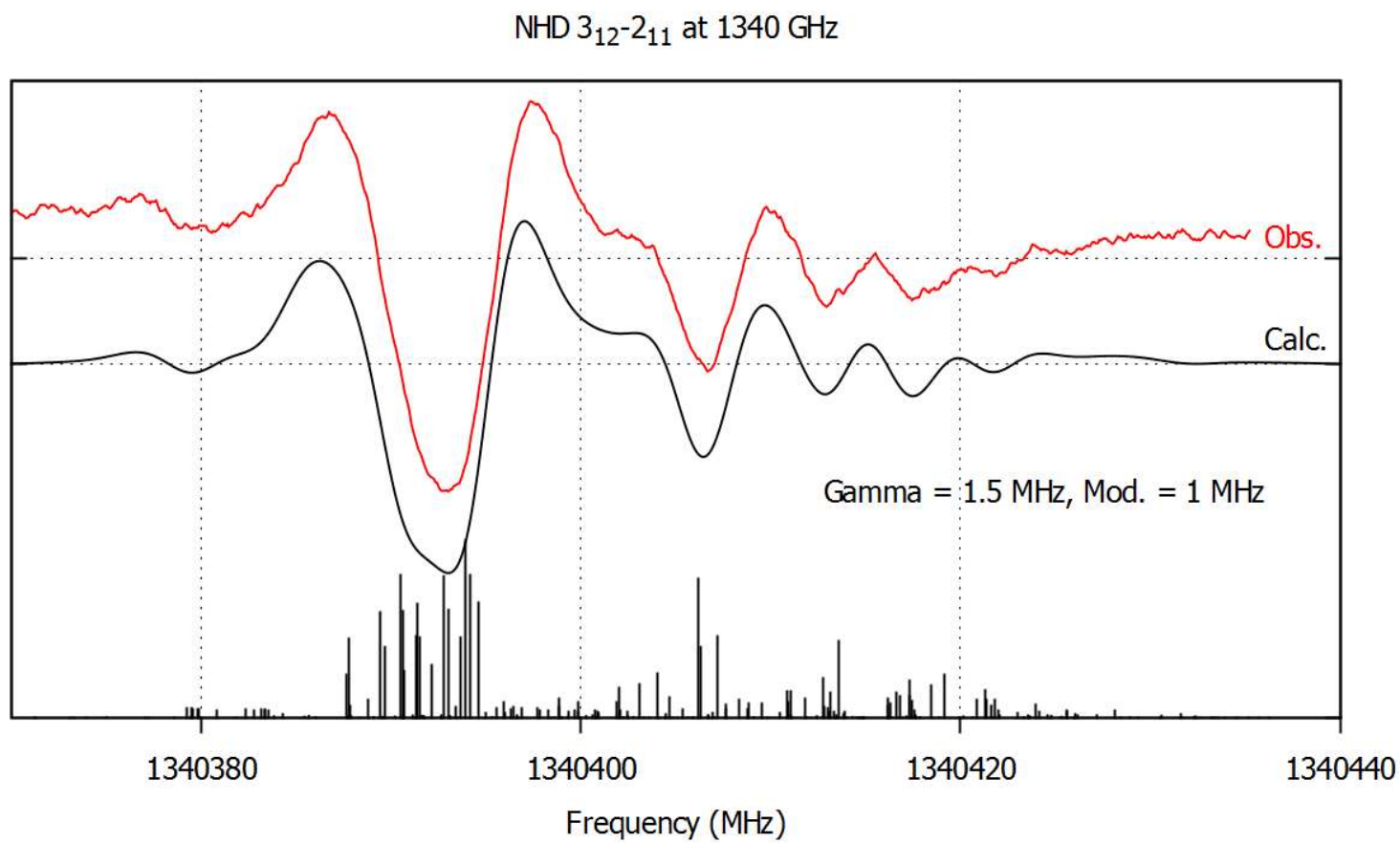




# Case 3 (NG)



# Case 4 (NG)



# Analysis of spectral profile obtained by frequency modulation technique

- Calculated spectral line profile  $S(\nu - \nu_0)$  is obtained by Fourier transformation of molecular polarization correlation function  $\Phi(t)$ , as

$$S(\nu - \nu_0) = I_o \cdot \exp \left[ -\Re \left( \mathcal{F} \left( J_2 \left[ mt \cdot \text{sinc} \left( \frac{\omega_m t}{2} \right) \right] \cos(\omega_m t) \cdot \Phi(t) \right) \right) \cdot L \right] \\ + \text{baseline correction}$$

- $\Phi(t)$  for each profile model is:

$$\text{Voigt} : \Phi_V(t) = \exp \left[ i\omega_0 t - \Gamma_0 t - \left( \frac{kv_{a0}t}{2} \right)^2 \right]$$

$$\text{Galatry} : \Phi_G(t) = \exp \left[ i\omega_0 t - \Gamma_0 t + \frac{1}{2} \left( \frac{kv_{a0}t}{\beta} \right)^2 \cdot \{1 - \beta t - \exp(-\beta t)\} \right]$$

$$\text{Speed Dependent Voigt} : \Phi_{SDV}(t) = \frac{\exp[i\omega_0 t - (\Gamma_0 - \frac{3}{2}\Gamma_2)t]}{(1 + \Gamma_2 t)^{3/2}} \cdot \exp \left[ -\frac{(kv_{a0}t)^2}{4(1 + \Gamma_2 t)} \right]$$



# NHD so far,,,

- 16 Rotational transitions are measured in the THz region (1.1 -1.6 THz)
- Some observed frequencies cannot be included in the fit due to complicated hyperfine structure
- Dedicated program for hyperfine structure envelope analysis must be developed

$\text{CHF}_2$  :

Di-fluoro methyl radical

Produce radical by MW disch. : Analysis of Tunneling motion

# Molecular structures of fluoromethyl radicals $\text{CH}_x\text{F}_{3-x}$ ( $x = 0 - 3$ )

- Vary from planar ( $x = 0$ ) to umbrella ( $x = 3$ )

$\tau$	$\text{CH}_3$	$\text{CH}_2\text{F}$	$\text{CHF}_2$	$\text{CF}_3$
MW / IR	$0^\circ$ <sup>1)</sup>	$\dot{=} 0^\circ$ <sup>2)</sup>	<b>This Work</b>	$18.15^\circ$ <sup>3)</sup>
ESR <sup>4)</sup>	$0^\circ$	$< 5^\circ$	$12.7^\circ$	$17.8^\circ$

<sup>1)</sup> C. Yamada et al., (1981)

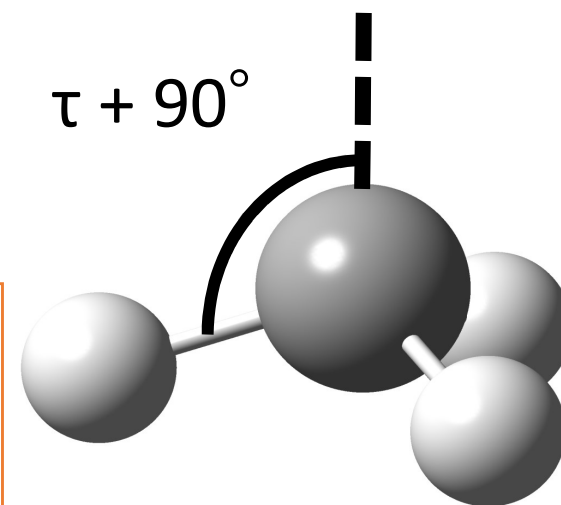
<sup>2)</sup> Y. Endo et al., (1983)

<sup>3)</sup> Y. Endo et al., (1982)

<sup>4)</sup> Fessenden & Shuler (1965)

“We may then expect that either or both of the two “intermediate” species  $\text{CH}_2\text{F}$  and  **$\text{CHF}_2$**  show the effect of **inversion doubling** in its high-resolution spectra”

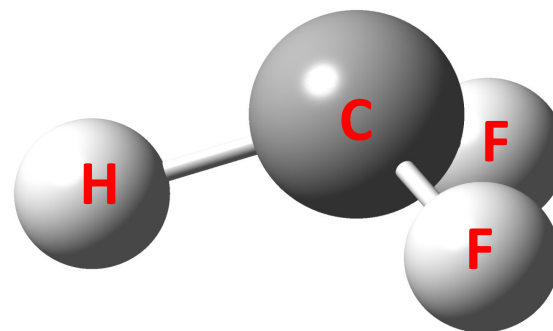
E. Hirota ,in “*High-resolution Spectroscopy of Transient Molecules* “



Definition of out-of-plane angle  $\tau$

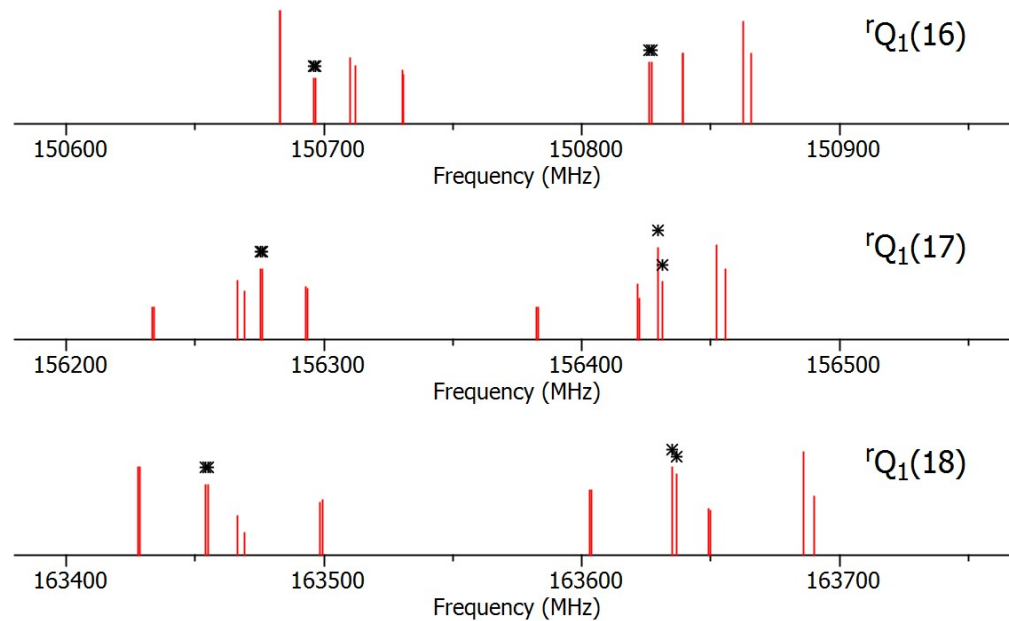
# Spectral Observation

- Radical produced by  
F(MW-disch. of  $\text{CF}_4$ ) +  $\text{CH}_2\text{F}_2$
- Spectrum observed in 90 – 400 GHz  
region and assigned 32 *b*-type rotational  
transitions (below 200 GHz)



# Observed spectra & Initial Analysis

- 4 spin system :  $S+I(F_{\text{tot}})+I(H)$ ,  $I(F_{\text{tot}})=I(F_1)+I(F_2)$
- Composite hyperfine structure of one H and two F nuclei giving  $I_F = 0$  (singlet or para) and  $I_F = 1$  (triplet or ortho)



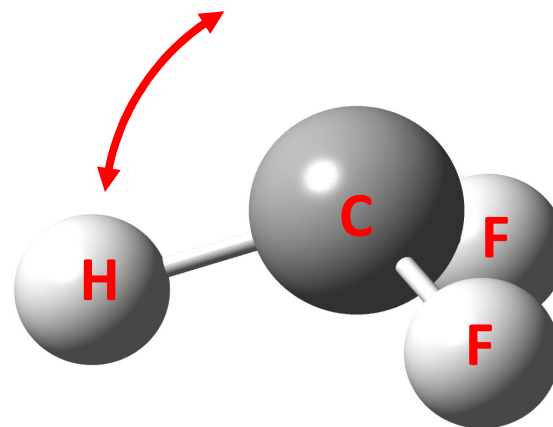
Stick diagram of observed hyperfine structure for the  $rQ_1$  branch transitions with  $N = 16 - 18$ .  
Asterisks indicate the hyperfine components of  $I_F = 0$

# Hyperfine structure of $\text{CHF}_2$

- 4 spin system :  $S + I(F_{\text{tot}}) + I(H)$ ,  $I(F_{\text{tot}}) = I(F_1) + I(F_2)$
- Composite hyperfine structure of one H and two F nuclei giving  
 $I_F = 0$  (singlet or para) and  $I_F = 1$  (triplet or ortho)
- Line frequencies of singlet changed **zigzag** relative to those of triplet

Hyperfine structure alternation  
in the inversion levels

Inada et al. CPL (1998)



# Conclusions

- Spectroscopy of short-lived species are interesting due to its diversity in many aspects such as production, analysis,,

From e-mail exchange more than 10 years ago,,,

Sounds like a lot of fun with lots of spins  $> 0$  !

*by Holger S. P. Müller*

I fully agree to you as many 'micorwavers' do !

*by Hiroyuki Ozeki*

# Collaborators

- Misaki Tanioka (Toho Univ.)
- Kaori Kobayashi (IMS, Toyama Univ.)
- Shuji Saito (IMS)
- Stephane Bailleux (Univ. Lille)
- Georges Wlodarczak (Univ. Lille)



JSPS KAKENHI

Grants-in-Aid for Scientific Research

Grant # 24540238, 15K05027, 18K03705



東邦大学