## PROBING THE ZEEMAN EFFECT IN LOW- $\Omega$ $F^4\Delta \leftarrow X^4\Delta$ TRANSITIONS IN FeH

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This work targets laboratory studies of the Zeeman effects in selected transitions of the FeH radical, observed in the atmospheres of dwarf stars. The  $F \leftarrow X$  electronic system falls around 1  $\mu$ m, and matches observation windows of the high-resolution spectropolarimeters SPIRou (brought into service in 2019) and ESPaDOnS mounted at the Canada-France-Hawaii Telescope. Many field-free line positions have already been reported for this radical from laboratory studies, notably from high-temperature sources<sup>ab</sup>.

Our earlier work with a sputter source<sup>c</sup> yielded information on magnetic response for just a few transitions between the two lowest spin components of the F and X states. We report here some preliminary results obtained from FeH formed in reaction between hydrogen atoms (generated in a microwave discharge of  $H_2$  in argon) and traces of iron pentacarbonyl vapour, at pressures around 1 Torr. This source<sup>d</sup> seems to produce more population in the  $X^4\Delta_{3/2}$  and  $X^4\Delta_{1/2}$  components of the ground state. Laser excitation of [1-0] transitions, with lock-in detection of fluorescence in the [1-1] band to eliminate laser scatter, has allowed some Zeeman-broadened profiles to be measured.

<sup>&</sup>lt;sup>a</sup>Line intensities and molecular opacities of the FeH  $F^4\Delta_i \leftarrow X^4\Delta_i$  transition; Dulick et al; Astrophys. J., <u>594</u>, 651-63, (2003)

<sup>&</sup>lt;sup>b</sup>The near-Infrared Spectrum of the FeH Molecule; Phillips et al; Astrophys. J. Supp. Ser., 65, 721-78, (1987)

<sup>&</sup>lt;sup>c</sup>Determination of Landé factors in the  $F^{4}\Delta_{\frac{5}{2},\frac{7}{2}}$  state of FeH by laser excitation spectroscopy; Crozet et~al.; J. Mol. Spectrosc.,  $\underline{303}$ , 46-53, (2014)

<sup>&</sup>lt;sup>d</sup>Detection of the free radicals FeH, CoH and NiH by far IR laser magnetic resonance; Beaton et al.; J. Chem. Phys. 89, 4446-48, (1988)