## COMPLEX ENVELOPE CHEMISTRY AND DYNAMICS OF NML CYGNUS

AMBESH PRATIK SINGH, Department of Molecular and Cellular Biology, University of Arizona, Tucson, AZ, USA; JESSICA L EDWARDS, LUCY M. ZIURYS, Department of Chemistry and Biochemistry, Department of Astronomy, The University of Arizona, Tucson, AZ, USA.

Oxygen-rich supergiant stars undergo intense, sporadic mass loss often causing them to have highly directed, non-isotropic outflows. These outflows have been previously identified in the envelope of the supergiant VY CMa, both in the infrared and in molecular lines. In VY CMa, these flows are best traced by sulfur-bearing molecules SO<sub>2</sub> and SO. To further investigate the unusual properties of supergiant stars, we have conducted a 1 mm survey of another such object, NML Cyg. Two major asymmetric outflows have been identified in the envelope of NML Cyg in the spectra of SO and SO<sub>2</sub>, one blue-shifted and the other red-shifted. Spectral line analysis with the radiative transport code ESCAPADE suggests that these two outflows are in random directions and are not bipolar in nature.