

CCH AND HNC IN PLANETARY NEBULAE

DEBORAH SCHMIDT, *Department of Astronomy, University of Arizona, Tucson, AZ, USA*; LUCY ZI-
URY, *Department of Chemistry and Biochemistry, University of Arizona, Tucson, AZ, USA*.

A survey of CCH and HNC has been conducted towards a sample of ten planetary nebulae of varying ages using the Submillimeter Telescope (SMT) of the Arizona Radio Observatory (ARO) at 1 mm. The $N = 3 \rightarrow 2$ transition of CCH at 262 GHz and the $J = 3 \rightarrow 2$ line of HNC at 272 GHz were observed using the ALMA Band 6 receiver at the SMT. The molecules were detected in most of the sources where HCN and HCO^+ had been identified in a previous survey. Molecular abundances for CCH and HNC have been determined in these nebulae, as well as $[\text{HCN}]/[\text{HNC}]$ ratios. These observations further support the notion that the chemistry in planetary nebulae remains active despite the ultraviolet radiation field from the central white dwarf star.